

# Fast Radio Bursts

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NCTS-TCA Summer Student Program 2022 workshop, 5 July 2022

# Outline

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## 1. Introduction:

FRBs and dispersion measure

## 2. The origin of FRBs

2-1. Localization

2-2. host galaxies

2-3. number density of FRBs

2-4. FRB classification

## 3. Applications of FRBs

If time is allowed

3-1. missing baryon problem

3-2. testing general relativity

3-3. dark energy

3-4. cosmic reionization

## 4. A new telescope plan in Taiwan: BURSTT

# Optical

A wide-field optical image of a starry night sky. The background is dark, filled with numerous small, bright stars of varying colors and magnitudes. In the center of the frame, there is a faint, diffuse, and elongated structure, possibly a nebula or a galaxy, with a slightly brighter core. The overall appearance is that of a deep-sky observation.

Credit: the Murchison Widefield Array (MWA)

# ***Intro:*** Fast radio burst (FRB)

Extra-galactic burst in radio

>10 Nature papers  
since 2019

Duration (millisecond)  
Brightness  
(~Jy)

Credit: Danielle Futselaar



FAST



CHIME

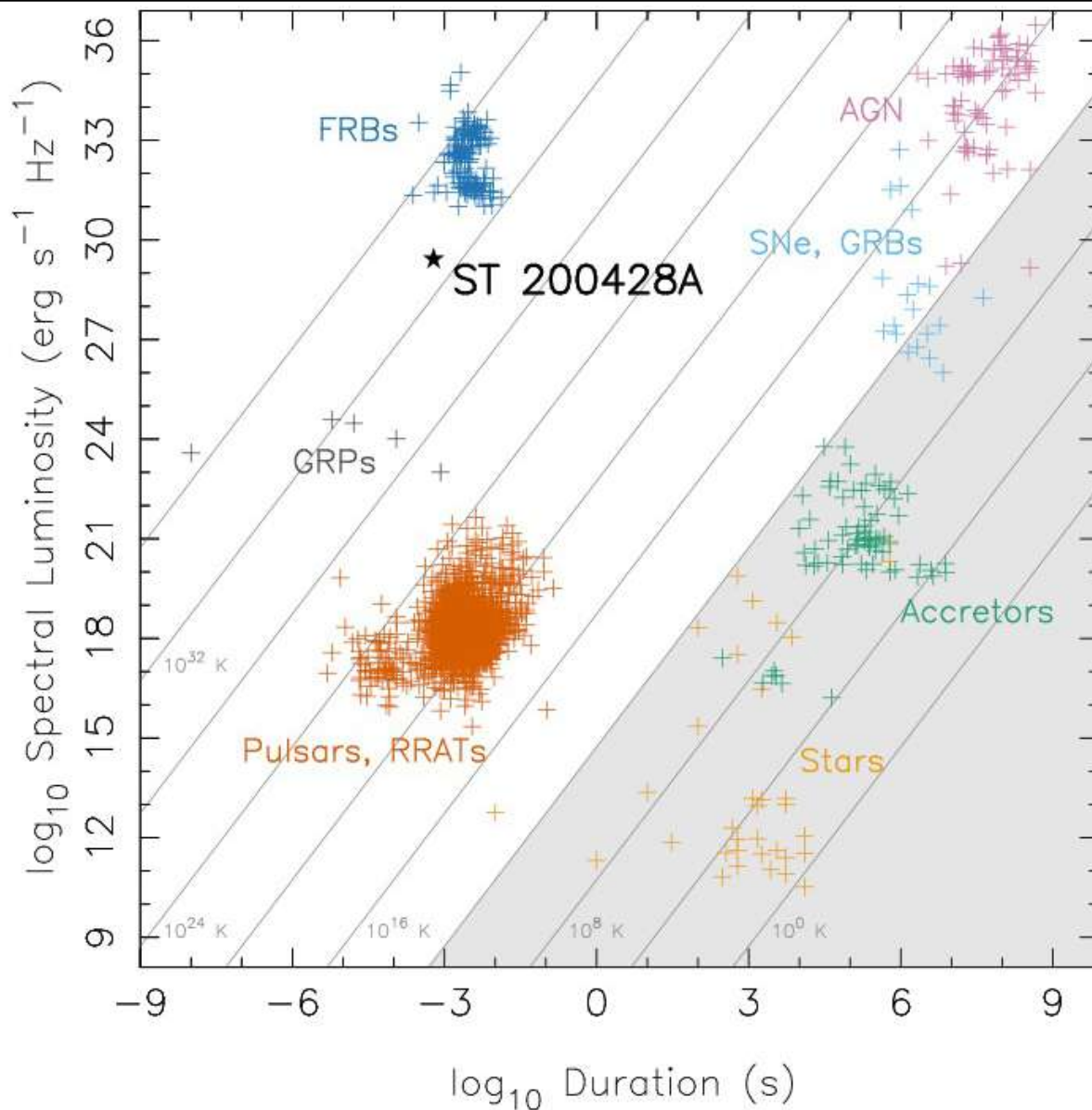


ASKAP



Parkes

# Intro: new parameter space



Bochenek  
et al 2020

# Game changers!

1000

>1000



FRBs per year

600

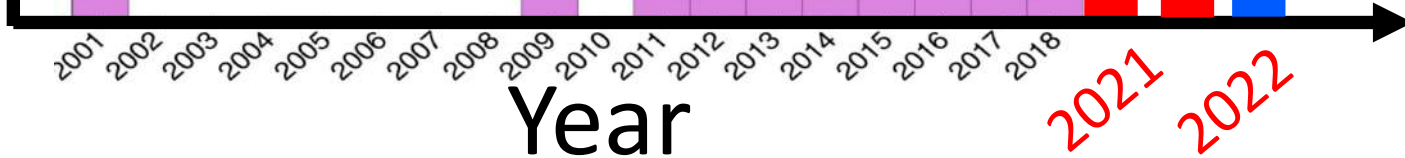
CHIME

FAST

20

- Parkes
  - UTMOST
  - GBT
  - Arecibo
  - ASKAP
- 61 FRBs in total

Keane 2018

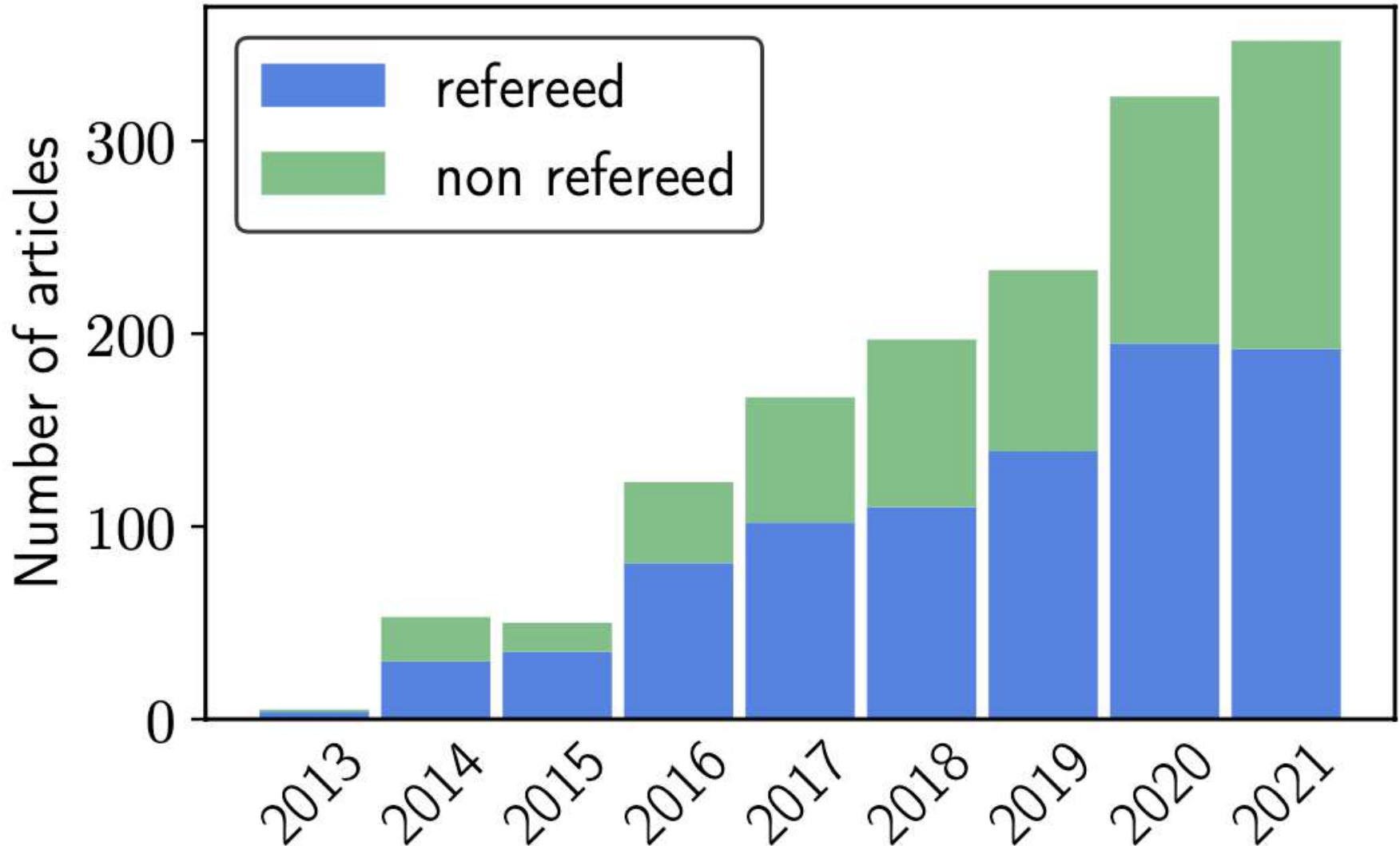


2021  
2022

# *Intro:* fast-growing field

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Petroff et al. 2021



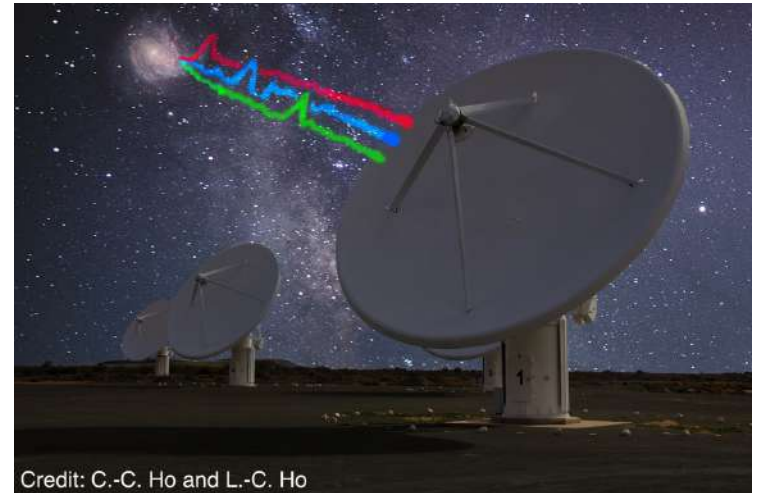
# *Intro:* Two types of FRBs

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Non-repeating FRBs



Repeating FRBs



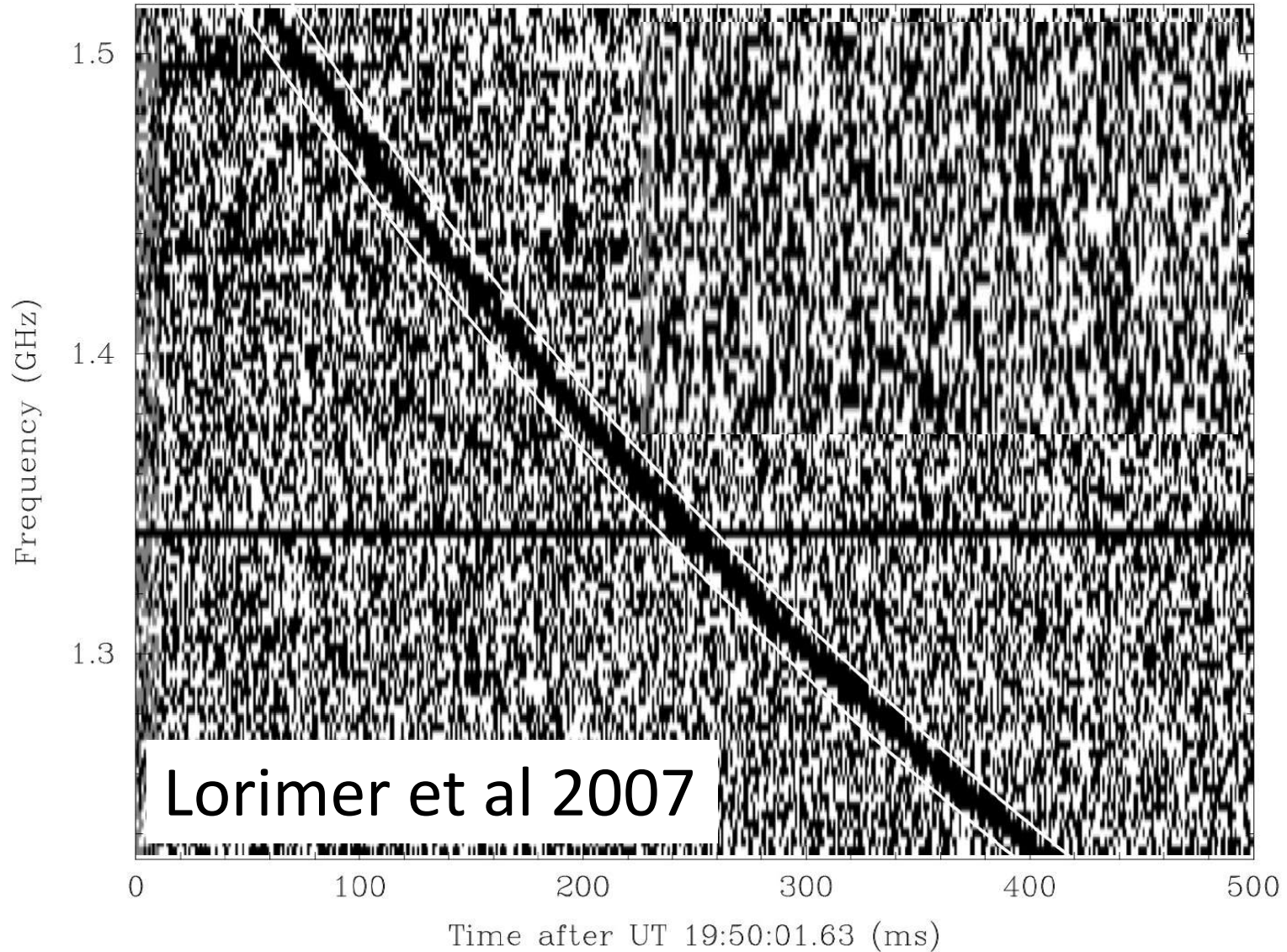
More than 1000 FRBs on sky every day!

Their origins are unknown



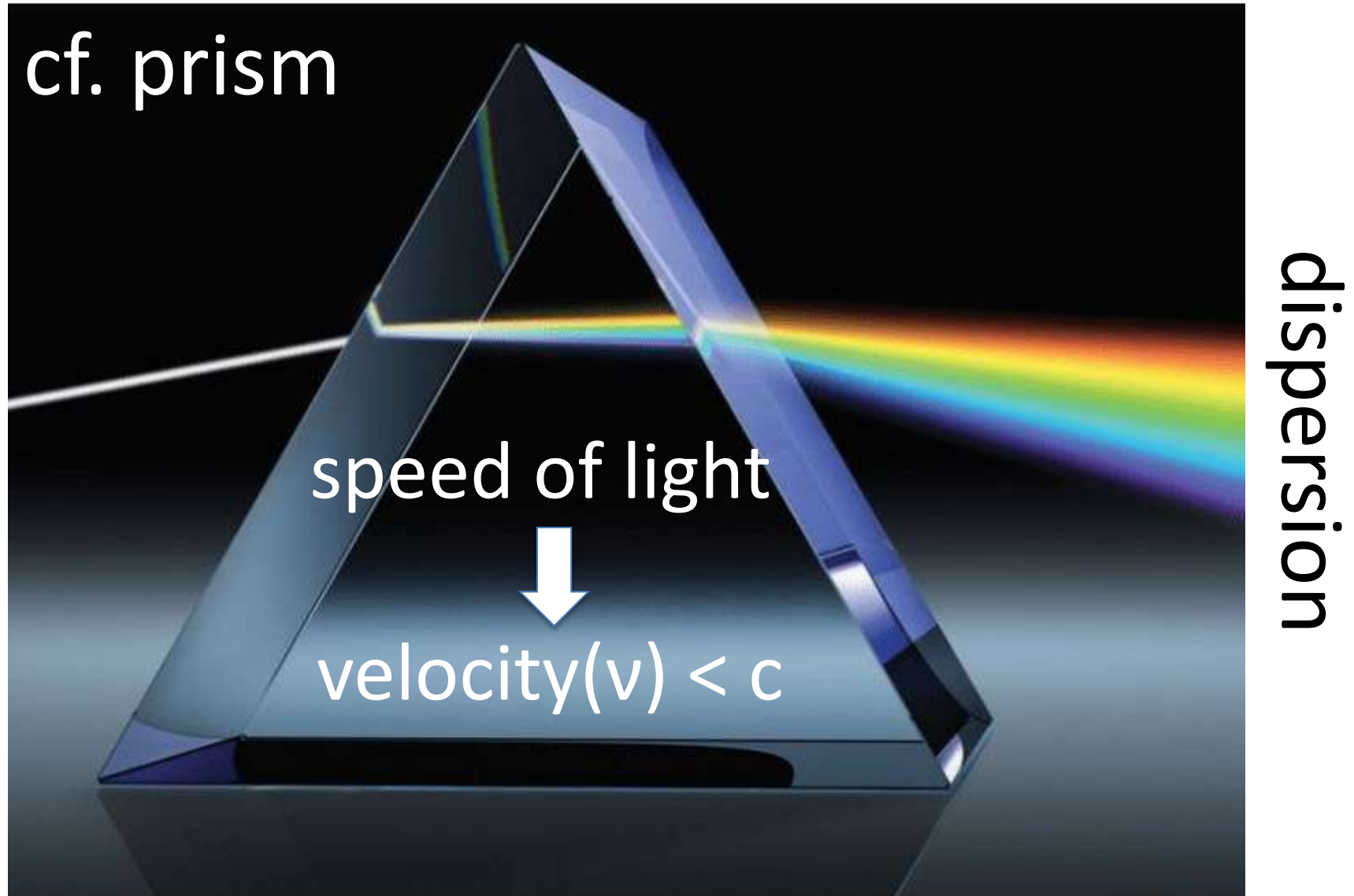
# *Intro:* the detection of an FRB

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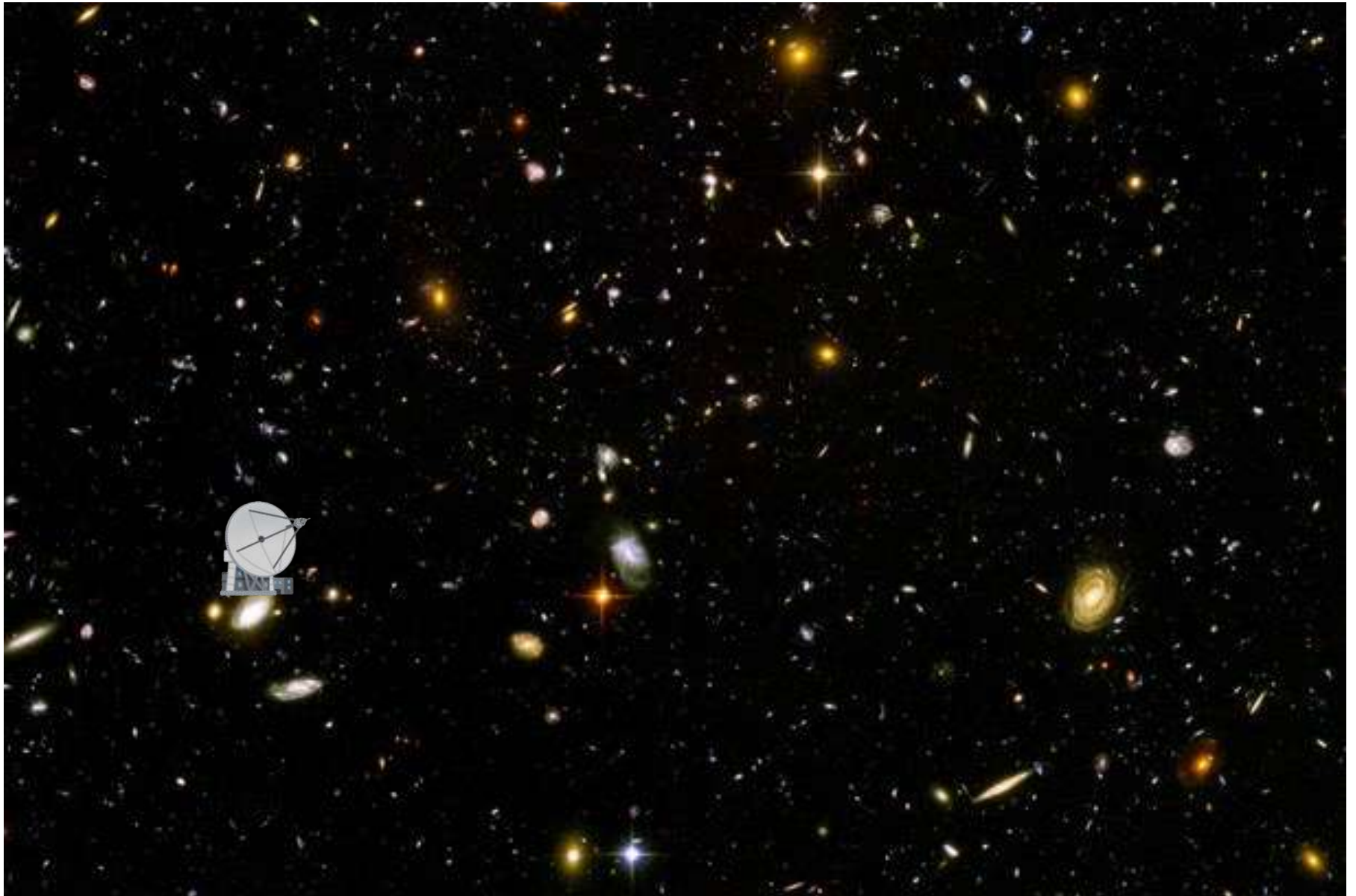


**Intro:** unique observable, 'dispersion measure'

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# ***Intro:*** the Universe is filled with plasma



# **Intro:** the speed of light changes in plasma

Speeds of radio emissions

high frequency: fast

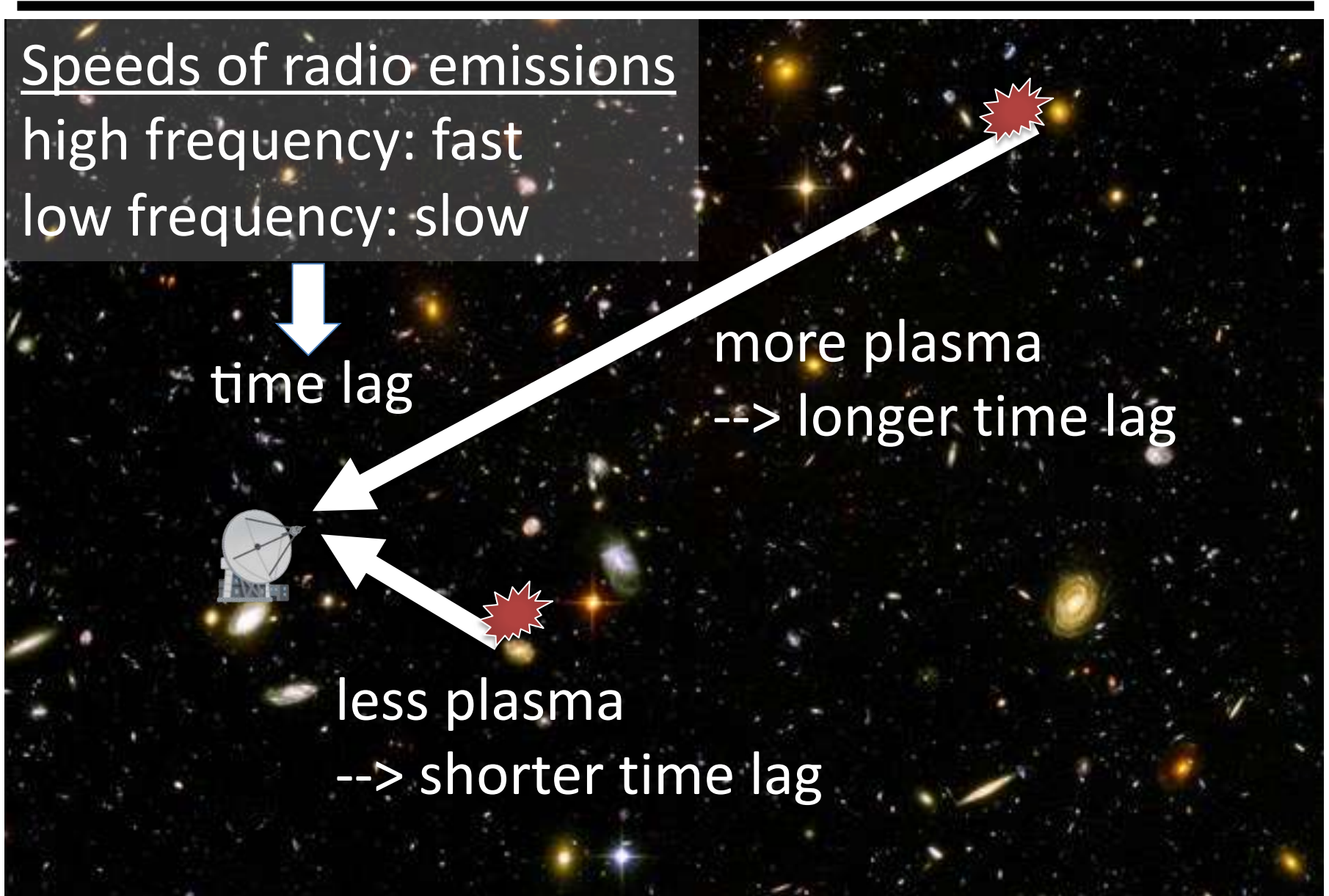
low frequency: slow

time lag

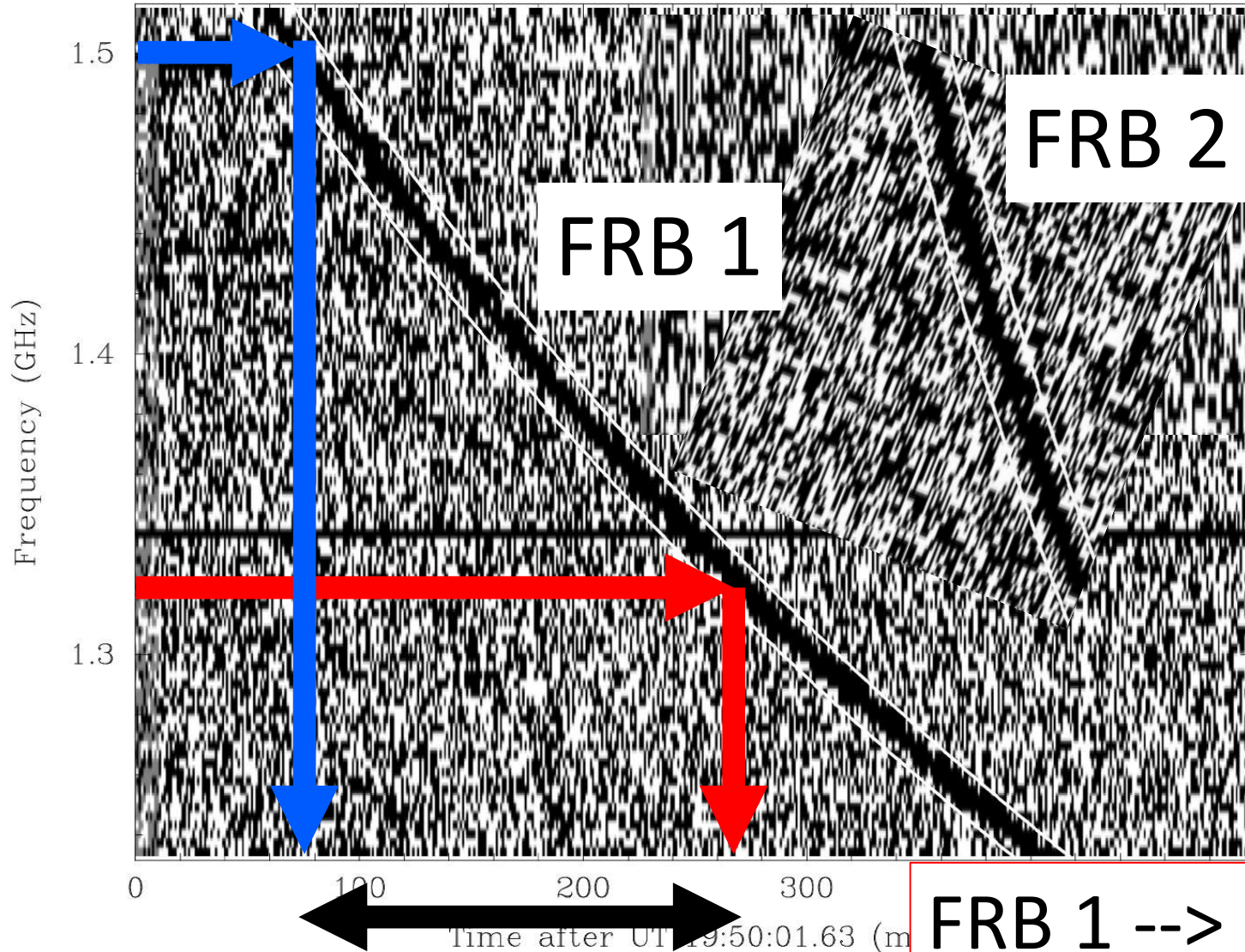


more plasma  
--> longer time lag

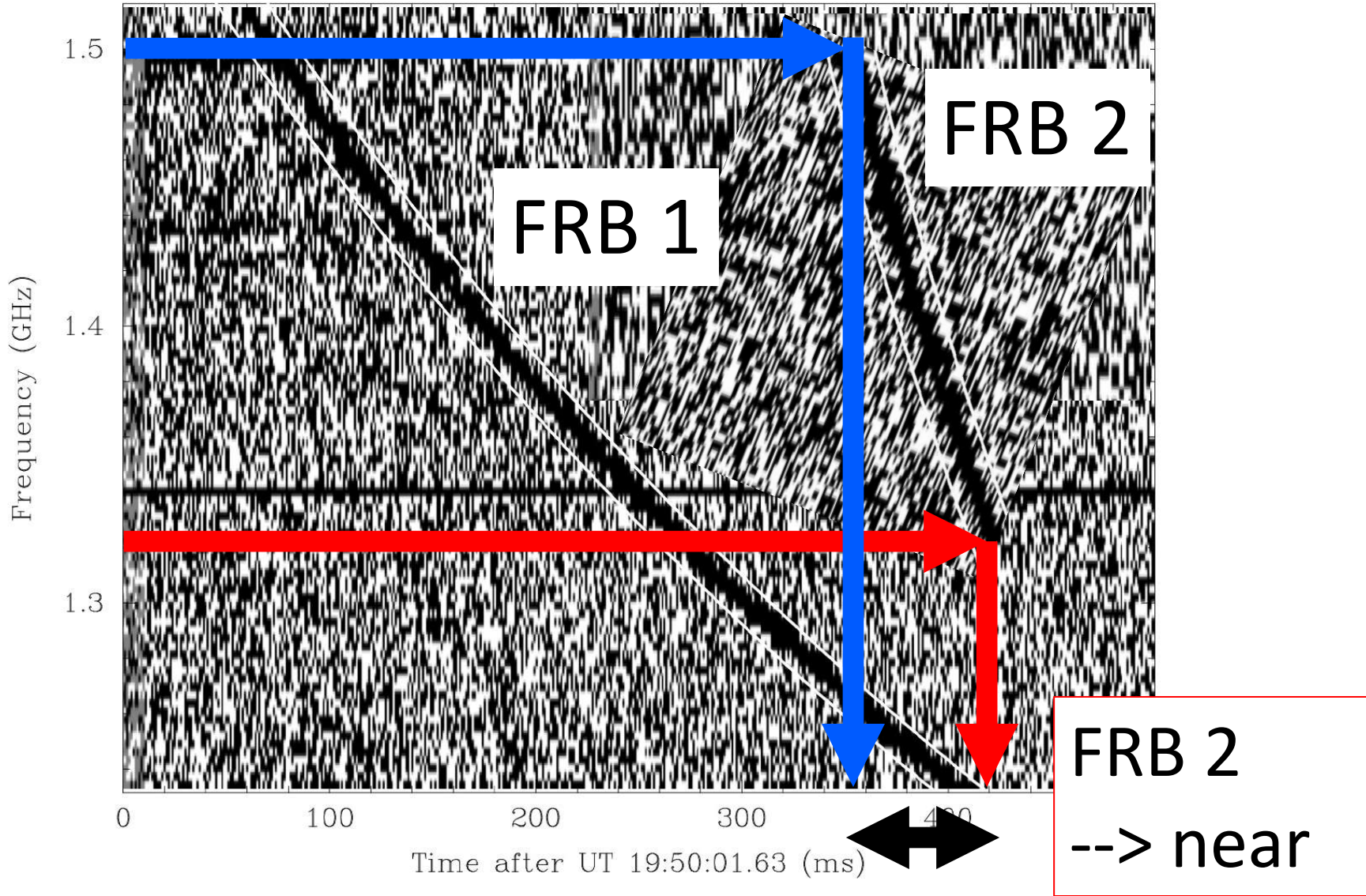
less plasma  
--> shorter time lag



**Intro:** time lag = dispersion measure  
= distance indicator



**Intro:** time lag = dispersion measure  
= distance indicator



## ***2. The origin of FRBs***

# *Origin:* possible FRB origins

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White dwarf



Old neutron star



Old stellar-mass black hole (BH)



Magnetar



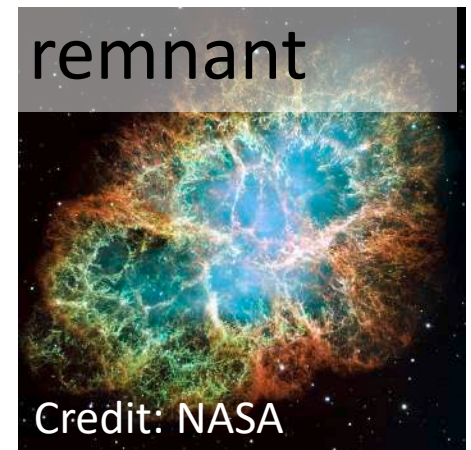
Young pulsar



Super massive BH

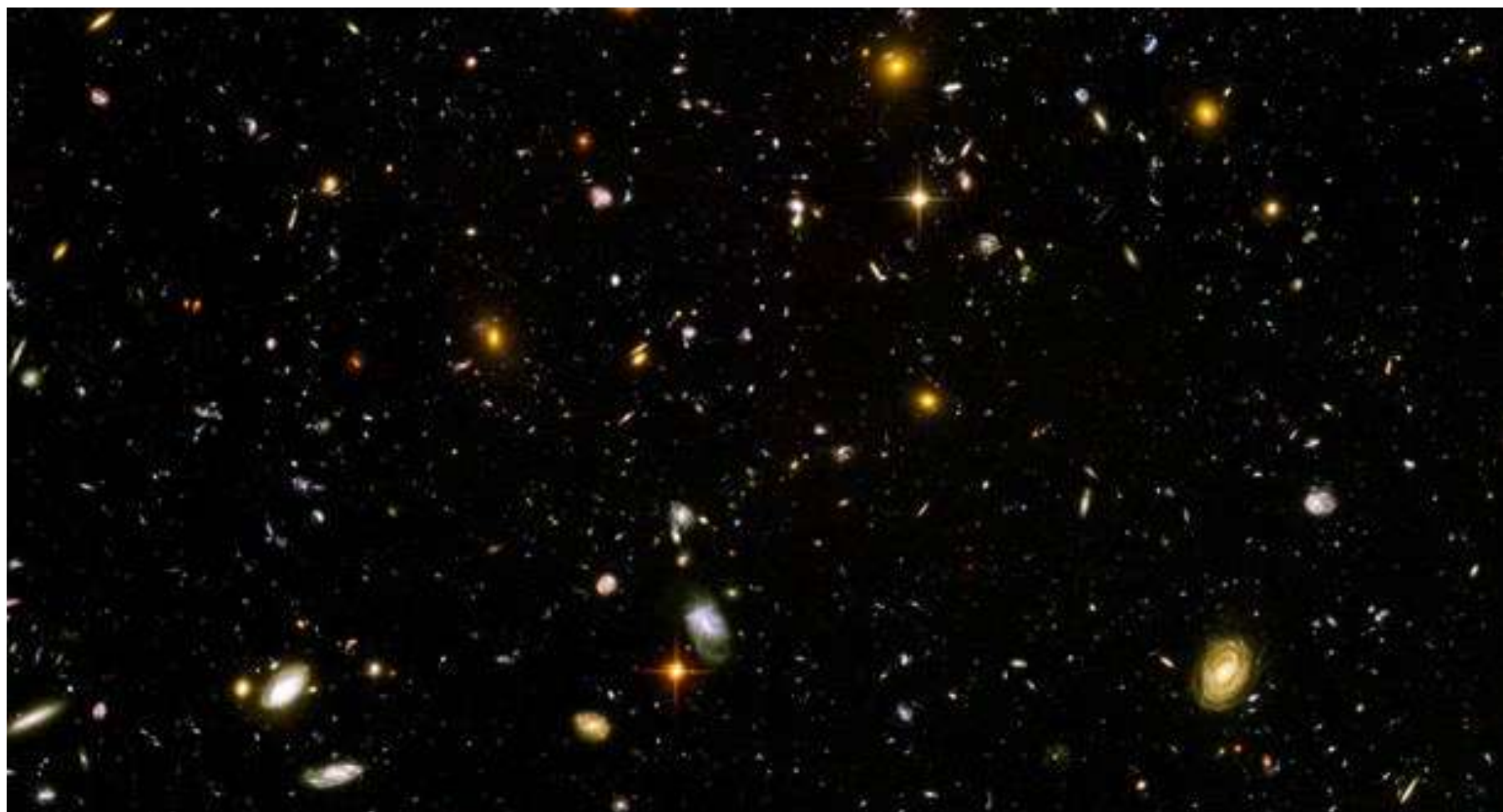


Supernova remnant





# *Origin:* How can we identify the progenitor?



Progenitor:  
astrophysical source which can emit FRB(s)

***Origin:*** where does the FRB come from?

Positional uncertainty (cf. CHIME)

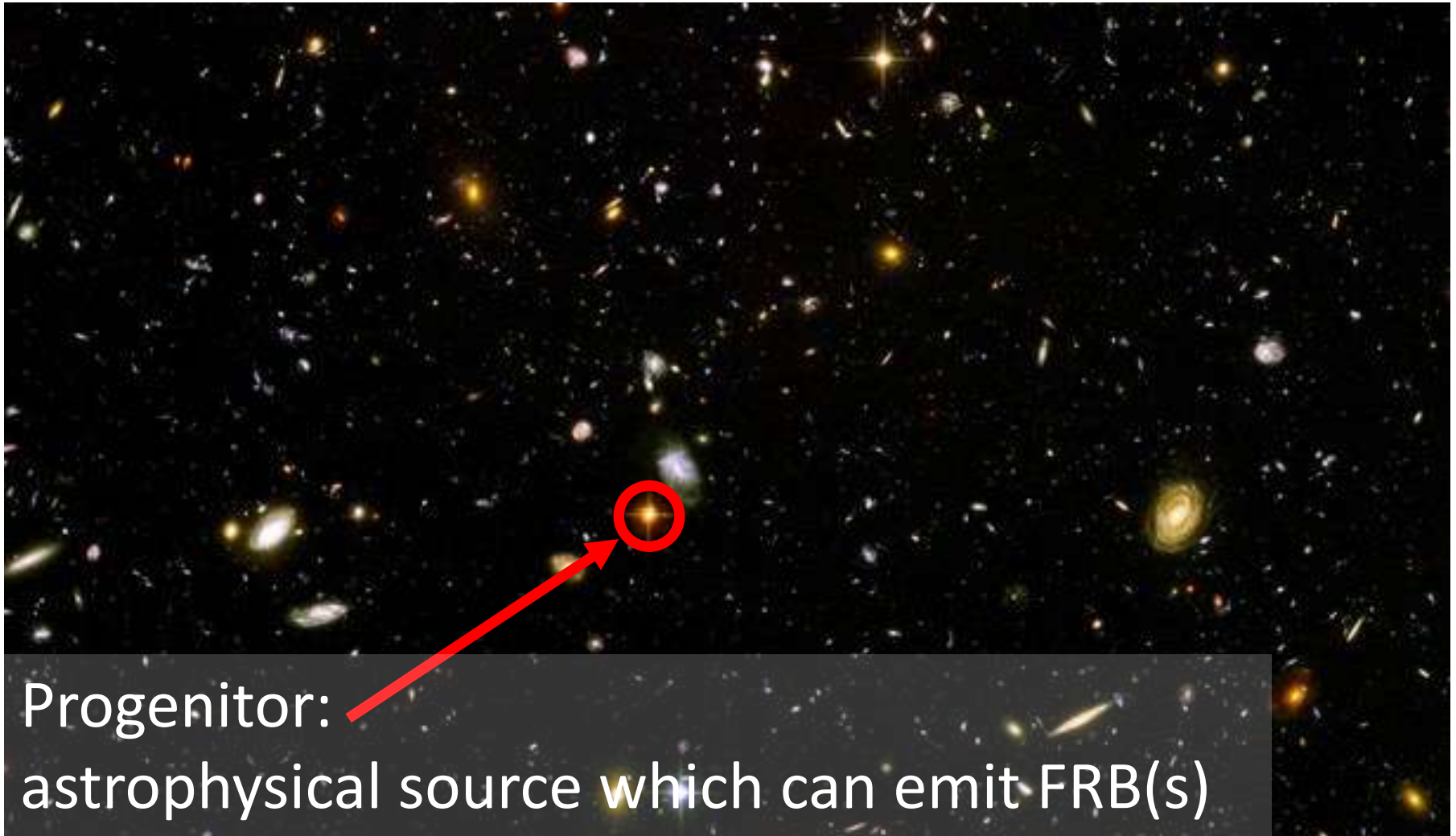


Progenitor:  
astrophysical source which can emit FRB(s)

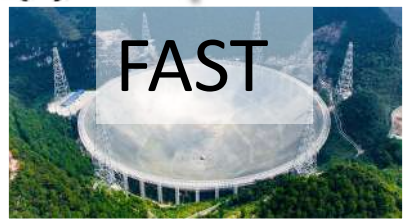
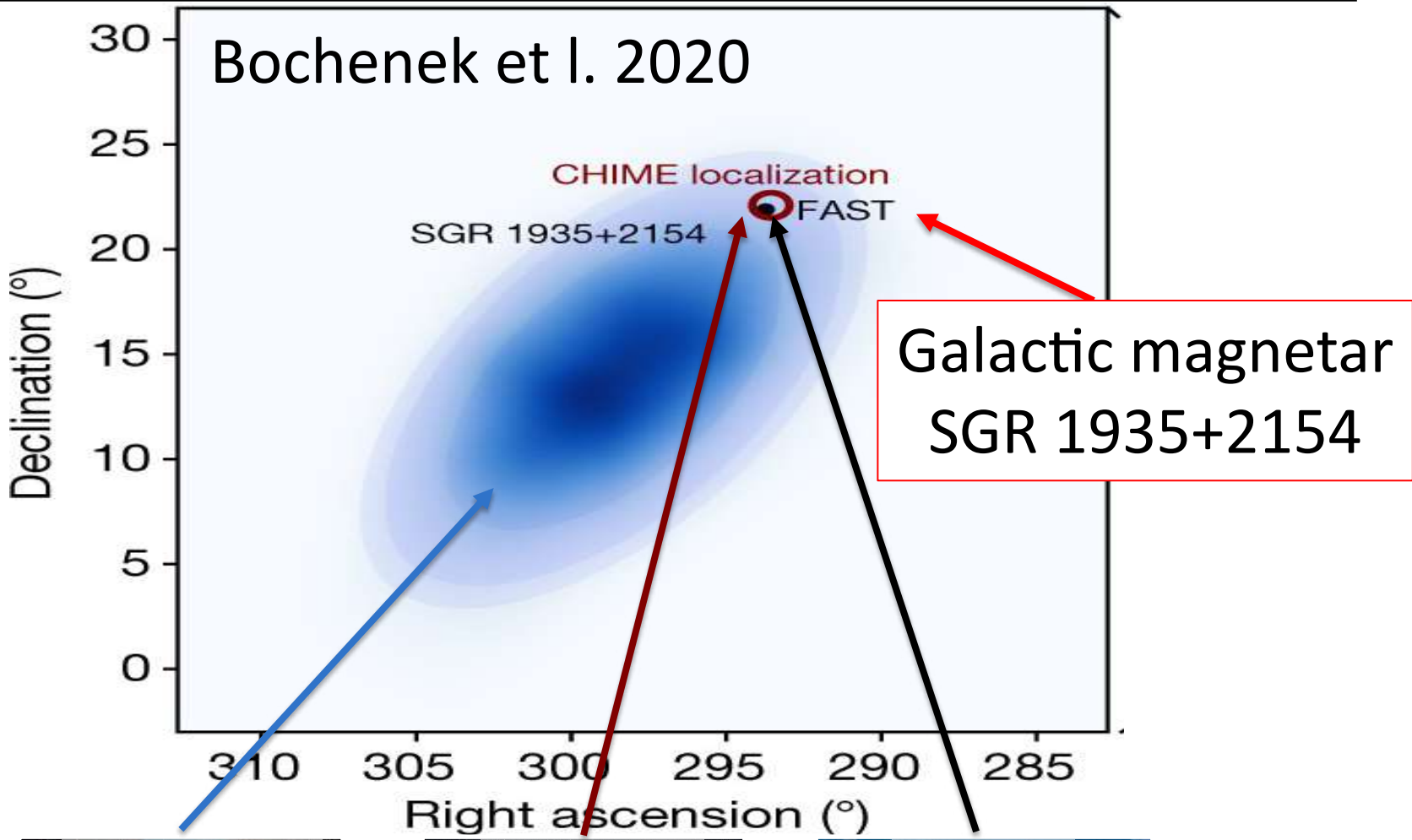
**Origin:** the localization identifies the progenitor

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Localization (cf. interferometric obs.)

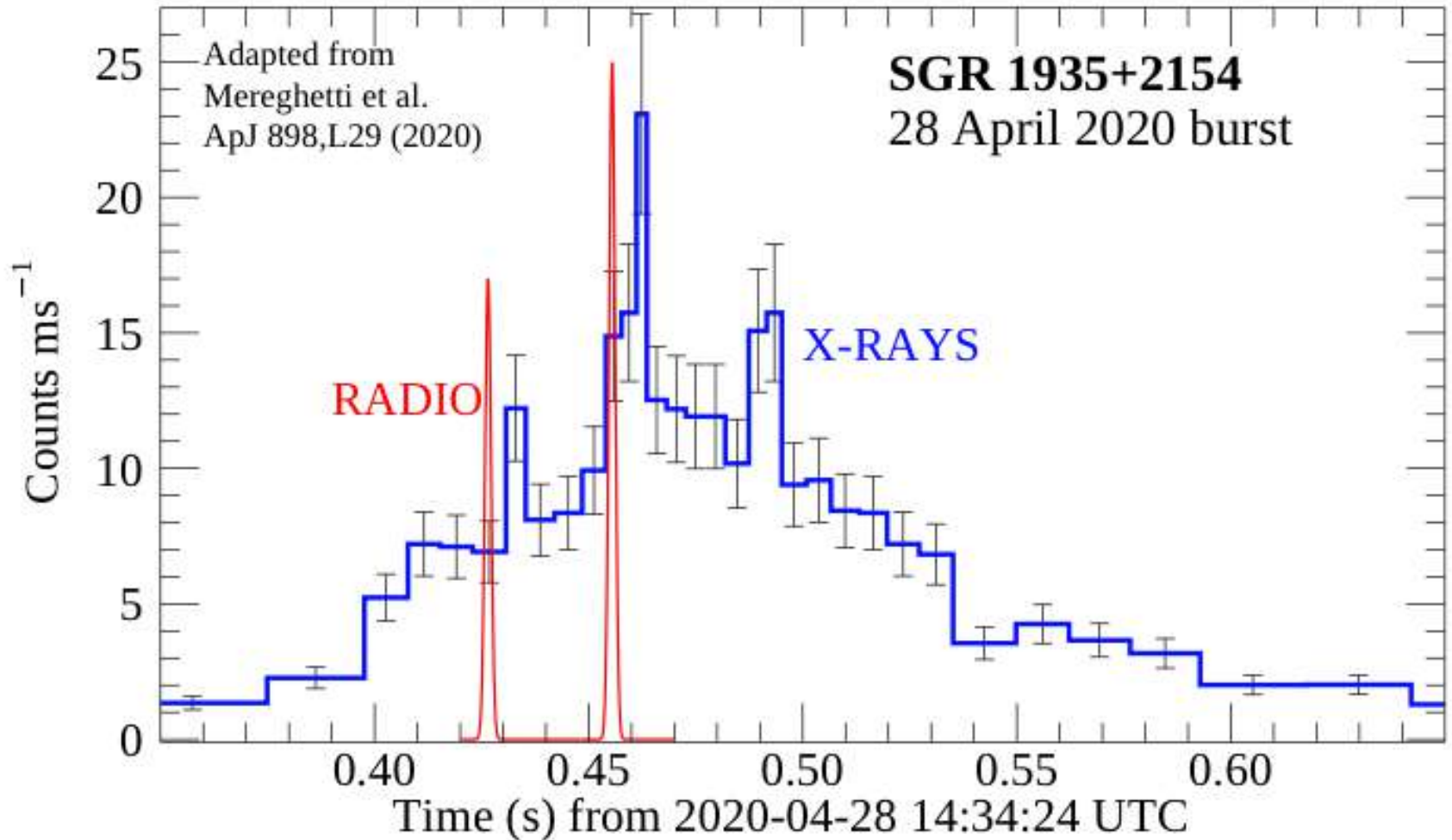


# Origin: the localized FRB at a Galactic magnetar



# *Origin:* the localized FRB at a Galactic magnetar

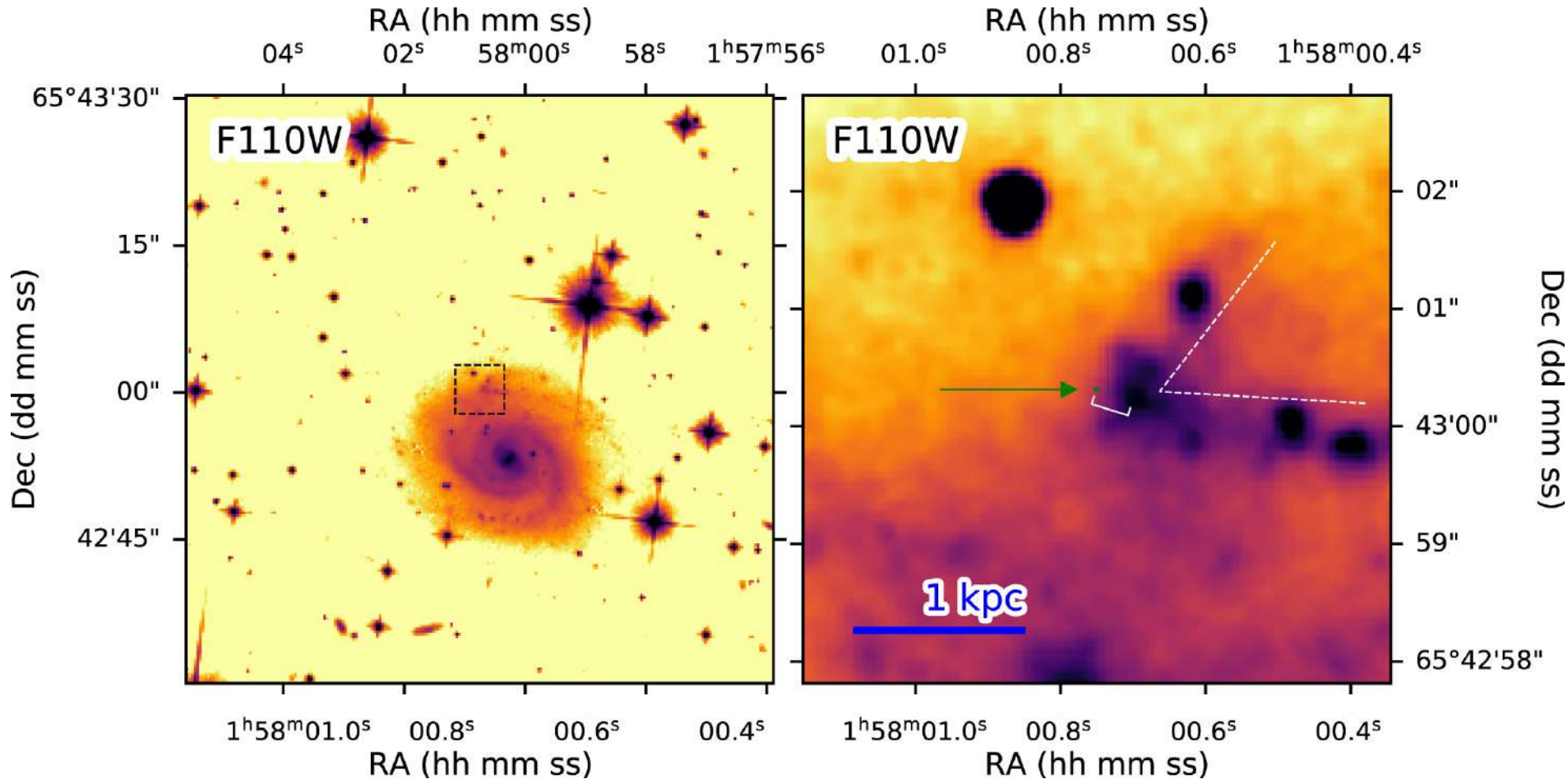
Mereghetti et al. 2020



Problem: no more 'direct' confirmation of the FRB progenitor

# *Origin*: an extragalactic repeating FRB source localized at a **star-forming region**

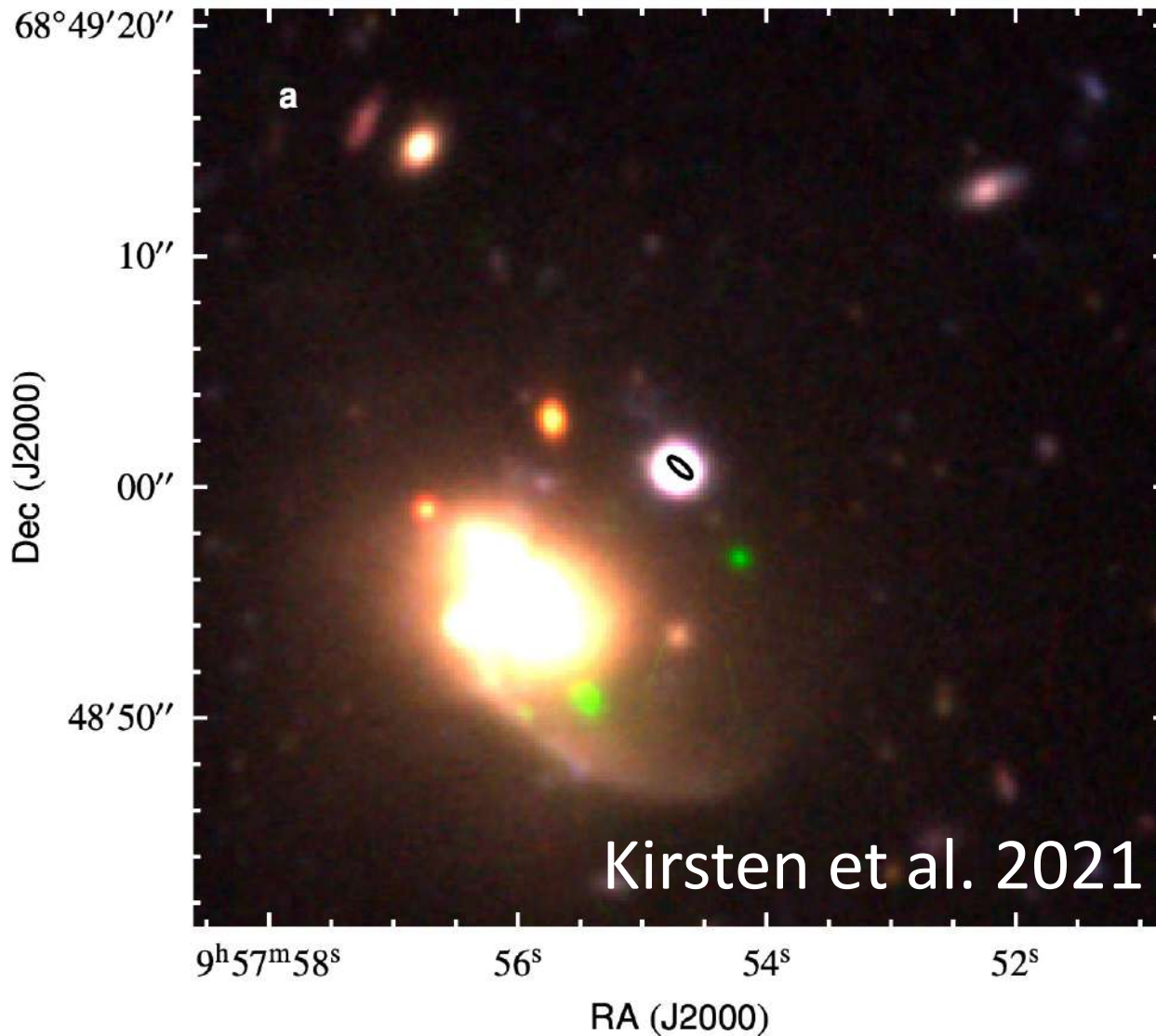
Tendulkar et al. 2021



FRB 20180916B --> star-forming region

***Origin:*** an extragalactic repeating FRB source  
localized at a **globular cluster**

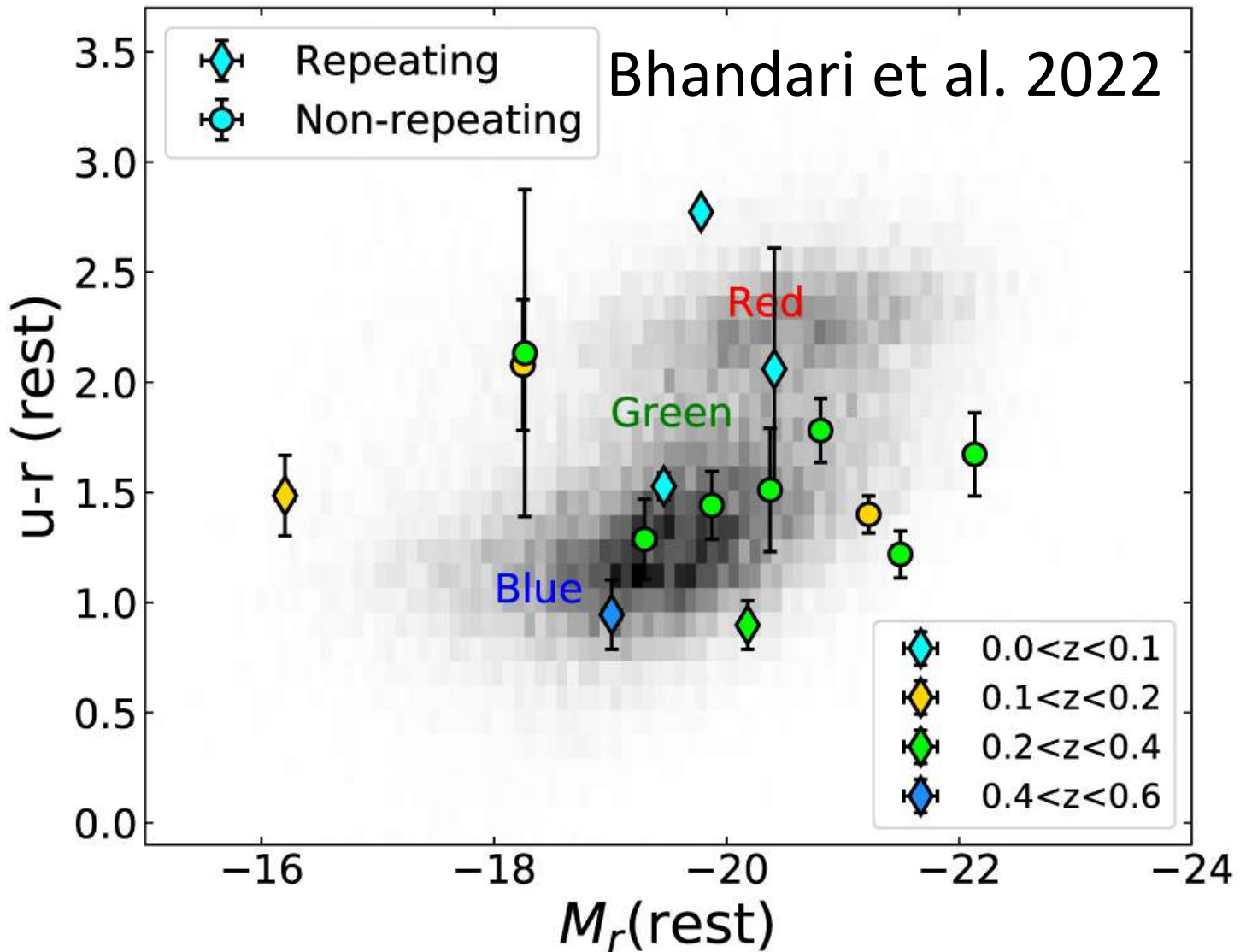
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FRB 20200120E --> globular cluster (old stellar system)

# *Origin:* FRB host galaxies:

## FRBs happen in any type of galaxies(?)



Problem: only ~20 host galaxies identified so far



# Problem

Previous research: tried to 'localize' FRB positions in the sky to reveal their origins  
→ didn't work well

# Solution (this work)

We changed the point of view  
→ focus on the history of FRBs  
(x10 more samples than before )



# *Origin:* possible FRB origins

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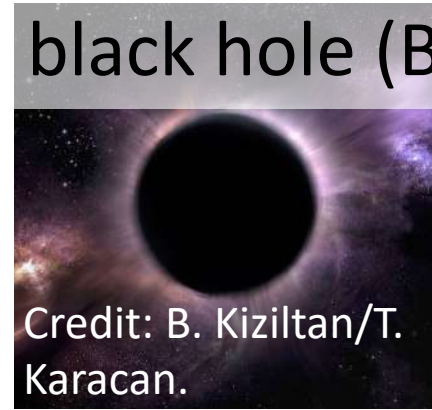
White dwarf



Old neutron star



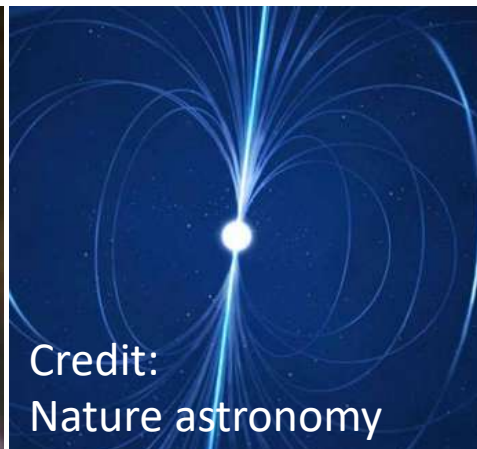
Old stellar-mass black hole (BH)



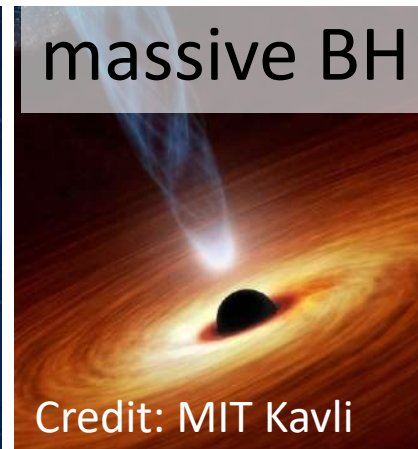
Magnetar



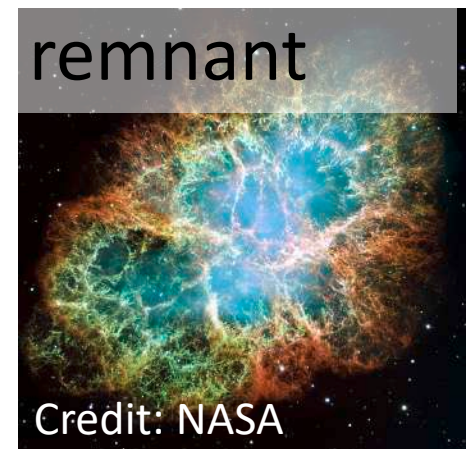
Young pulsar



Super massive BH



Supernova remnant



# *Origin:* possible FRB origins

White dwarf



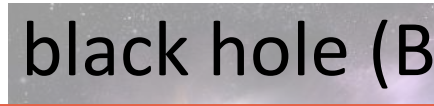
Credit: Tetsuya Hashimoto

Old neutron star



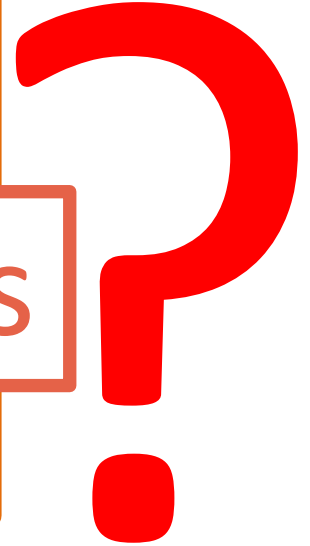
Credit: Mark Garlick

Old stellar-mass black hole (BH)



Credit: B. Kiziltan/I. Karacan.

Old objects  $\propto$  stellar mass



Magnetar



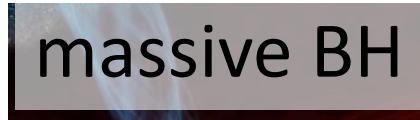
Credit: Tetsuya Hashimoto

Young pulsar



Credit: Nature astronomy

Super massive BH



Credit: MIT Kavli

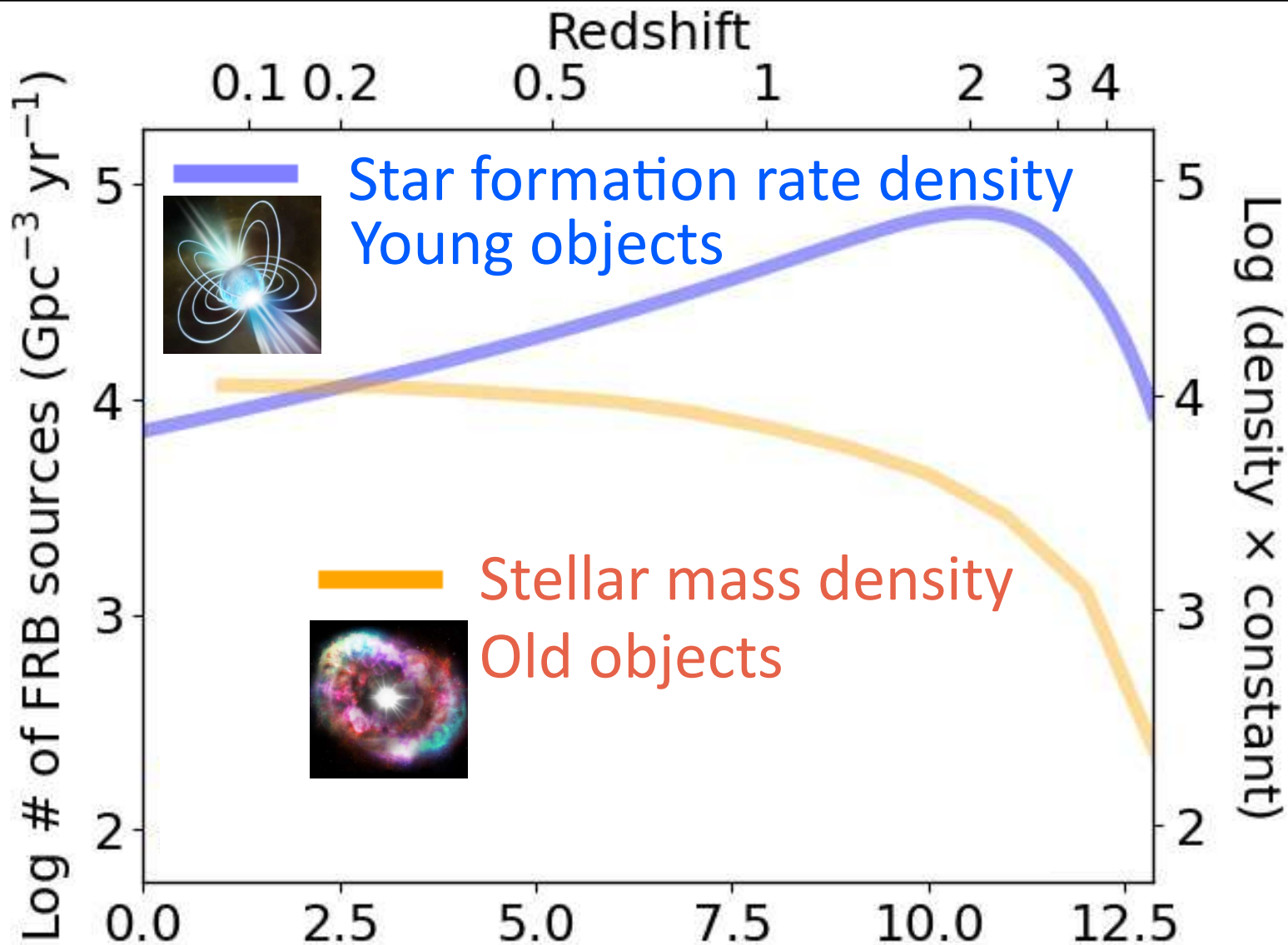
Supernova remnant



Credit: NASA

Young objects  $\propto$  star formation

# Origin: Old vs Young

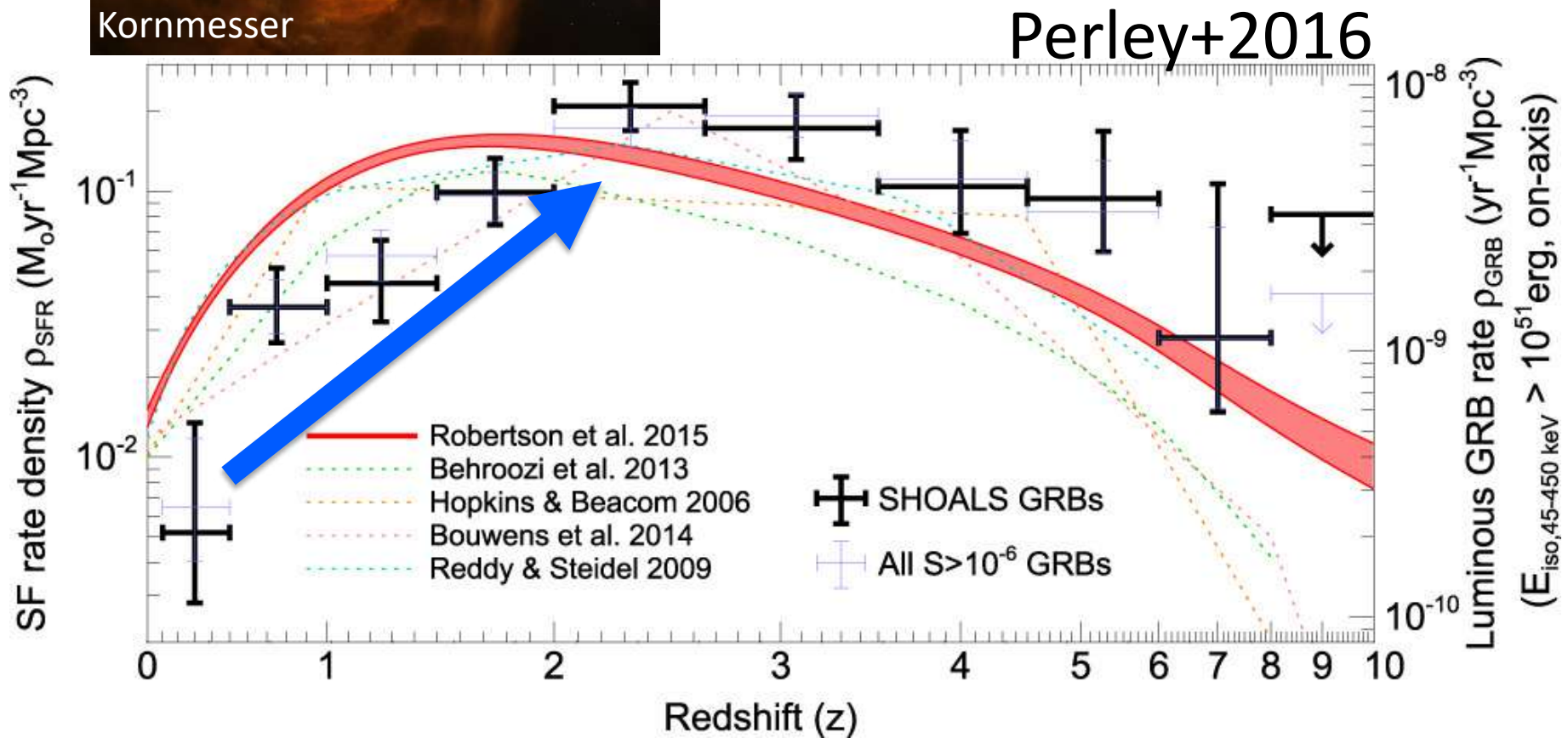


Nearby Universe      Lookback time (Gyr)      Distant Universe

# Origin: example (LGRBs)



Long Gamma-ray bursts  
 $\propto$  star formation



***Origin:*** Let's see the answer!

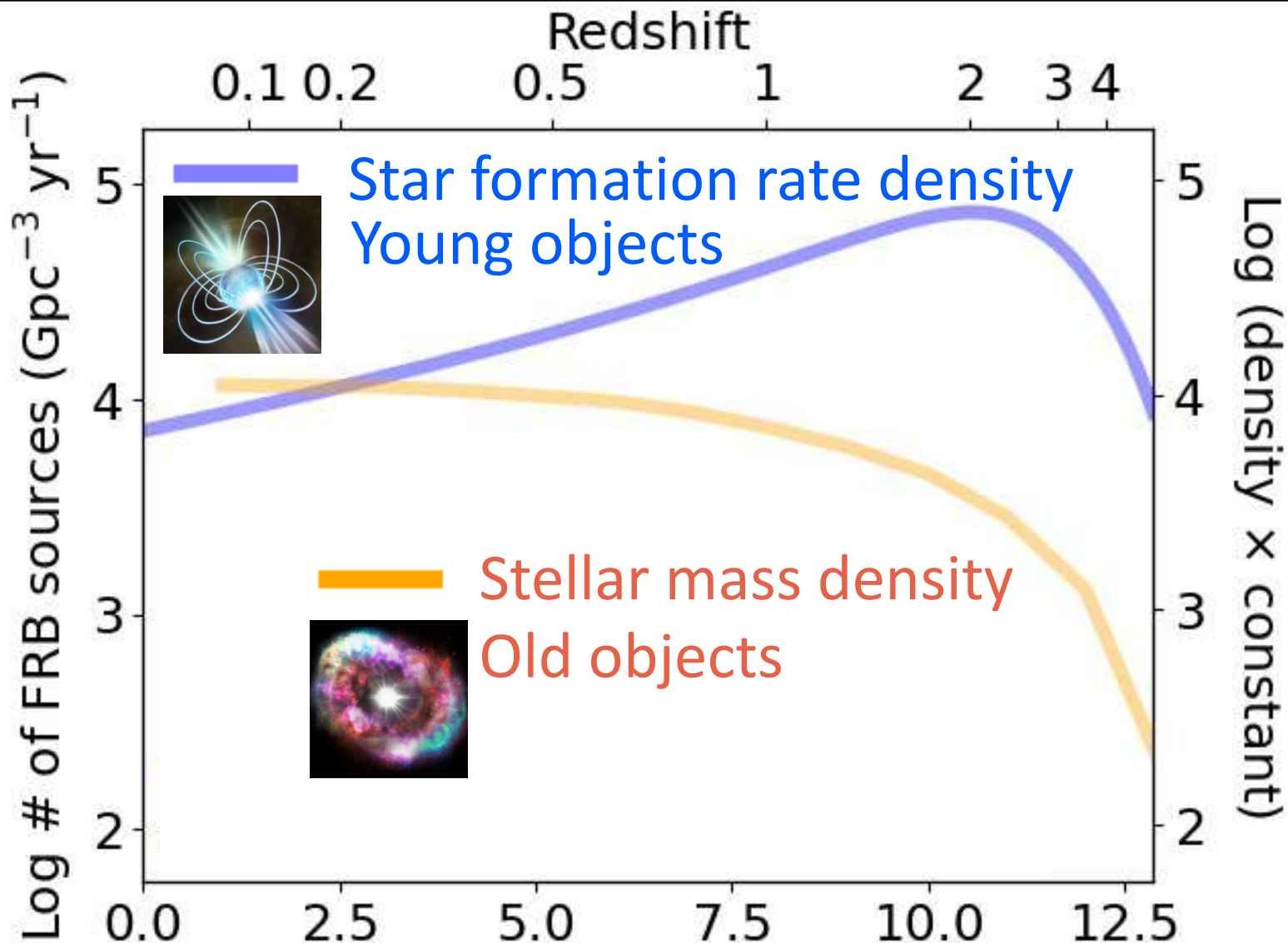
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Credit: CHIME

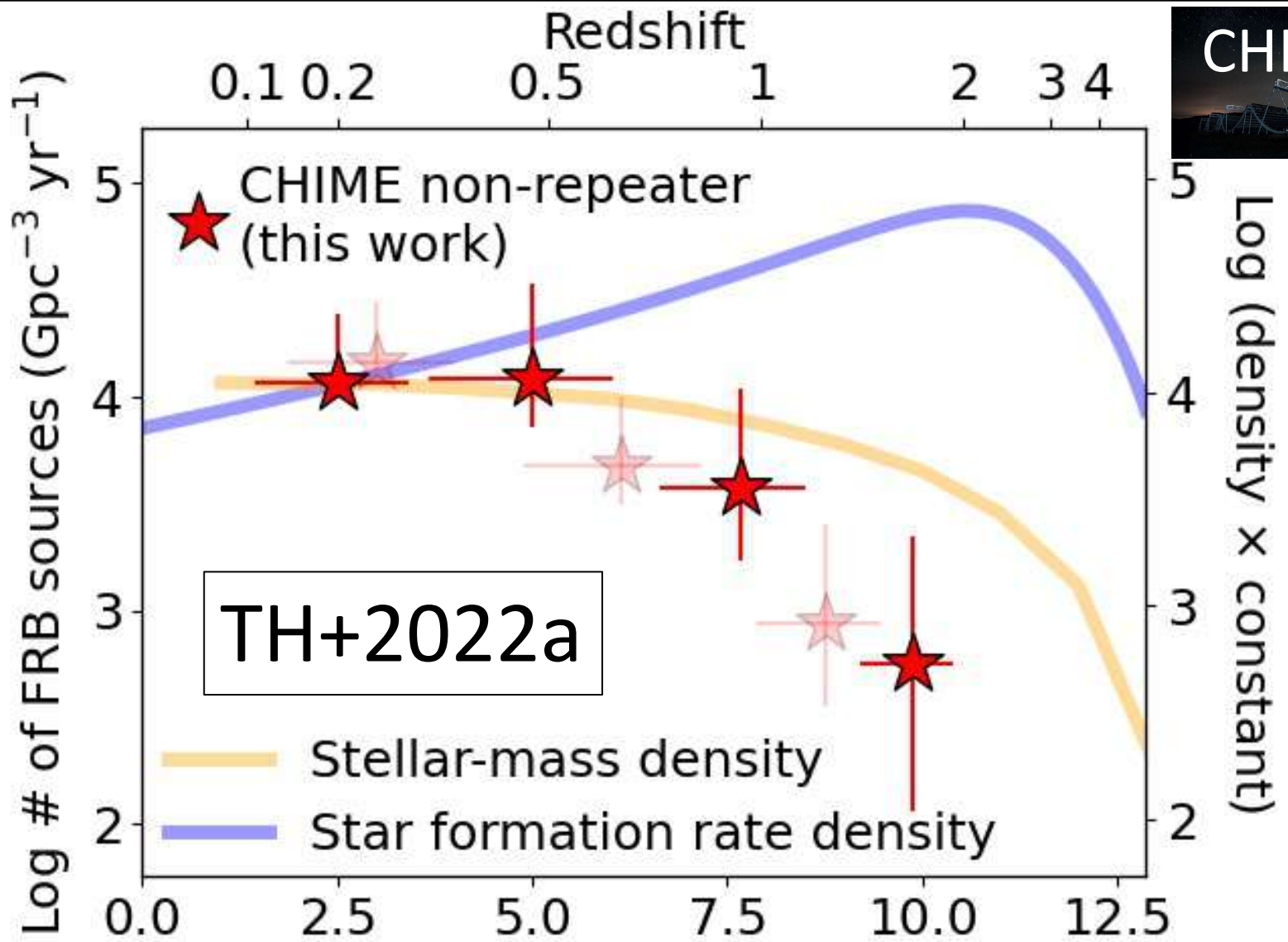
**CHIME: ~500 non-repeating FRBs in 2021!**

# Origin: Old vs Young



Nearby Universe      Lookback time (Gyr)      Distant Universe

# Our result: Non-repeater → Old





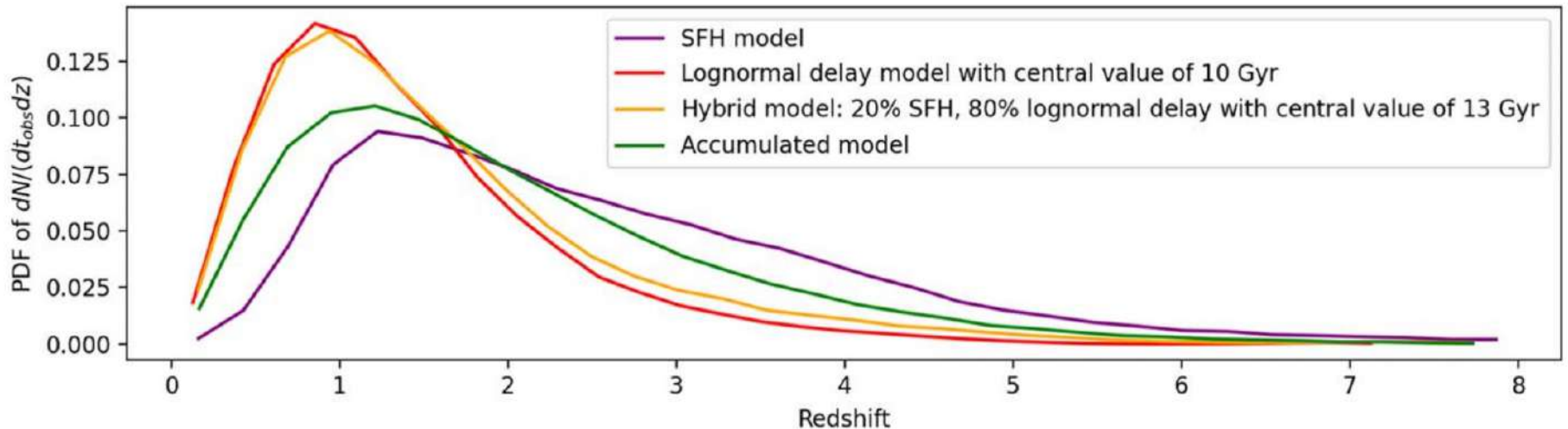
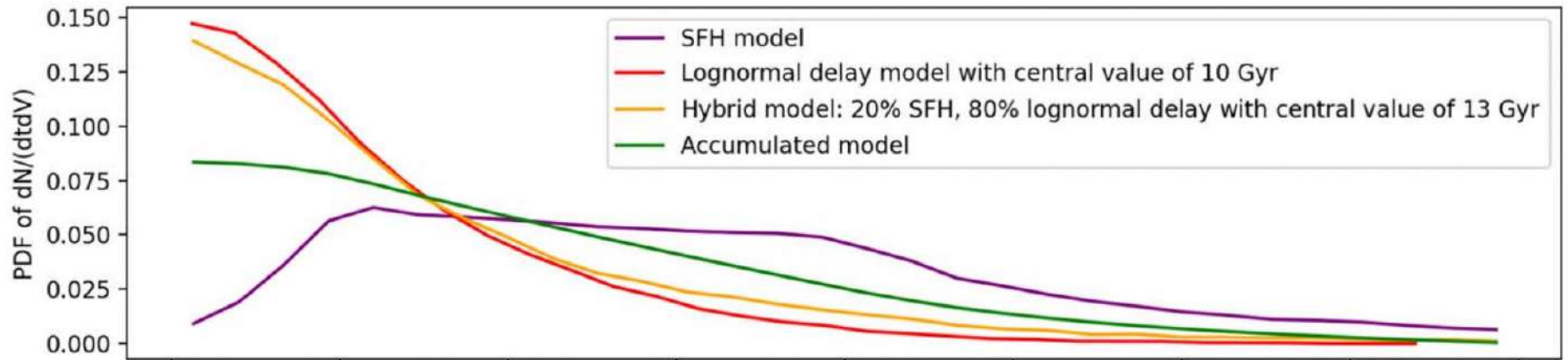
# *Origin:* other FRB population analyses

Author	free from the z-evolution assumption?	test an old population scenario?	sample	homogeneous sample?	conclusion
TH+2022a	Yes	Yes	CHIME	Yes	Old pop.
Zhang&Zhang 2022	No	Yes	CHIME	Yes	Old pop.
James+2022	No	No	ASKAP/ Parkes	No	Young pop.
Arcus+2021	No	Yes	ASKAP/ Parkes	No	Both young and no-evo pops.
Zhang+2021	No	Yes	ASKAP/ Parkes	No	Both pops.
TH+2020c	Yes	Yes	Parkes	Yes	Old. pop.

# Origin: FRB population model

$$dN/dE \propto E^{-\alpha}$$

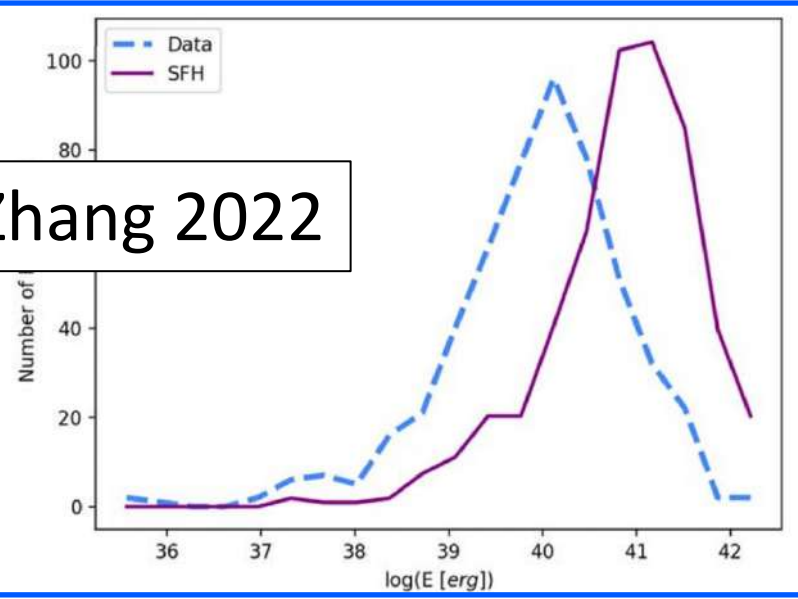
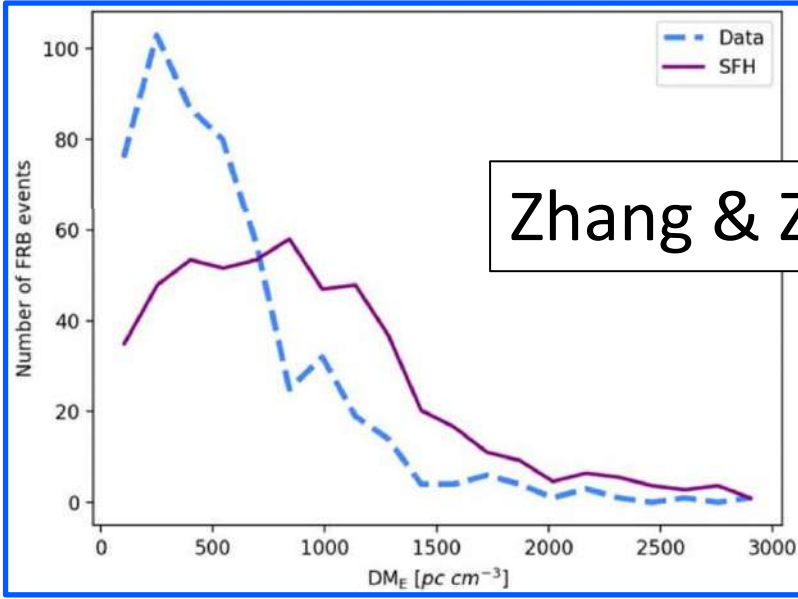
Zhang & Zhang 2022



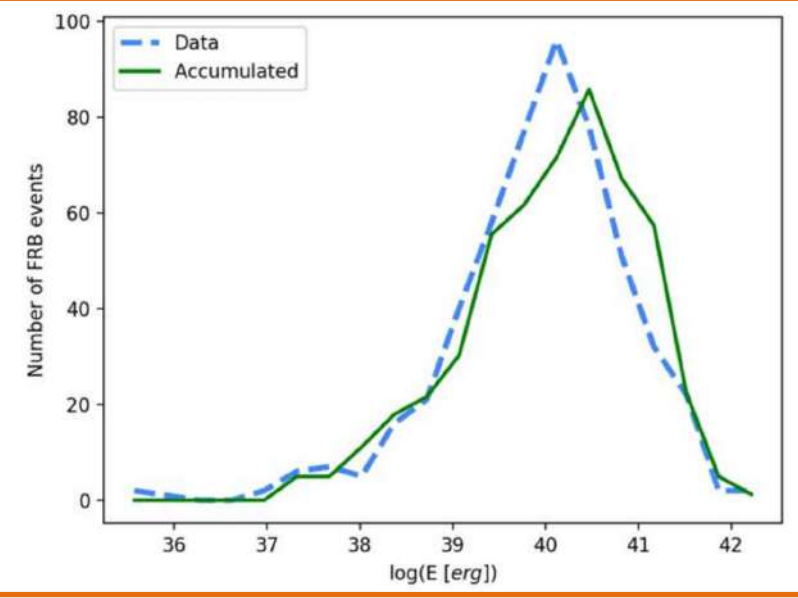
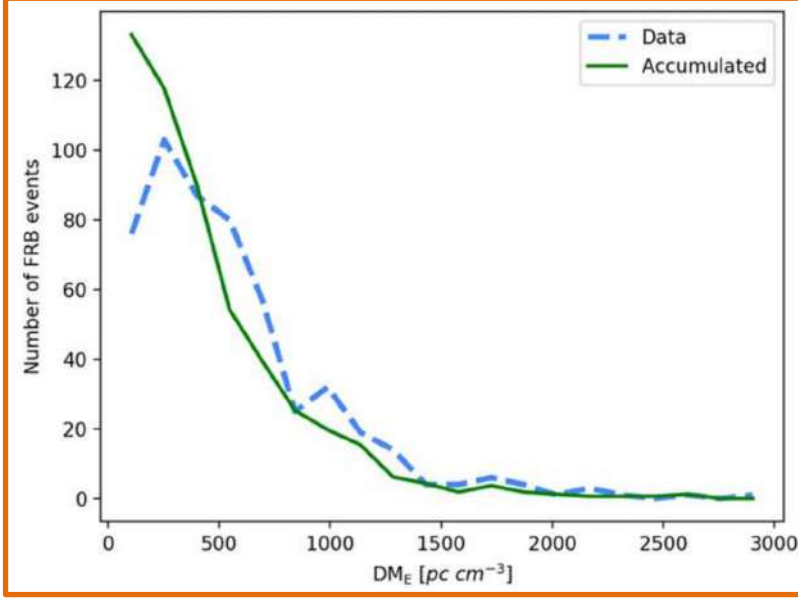
# Origin: the stellar-mass scenario better explains

Star formation

Zhang & Zhang 2022

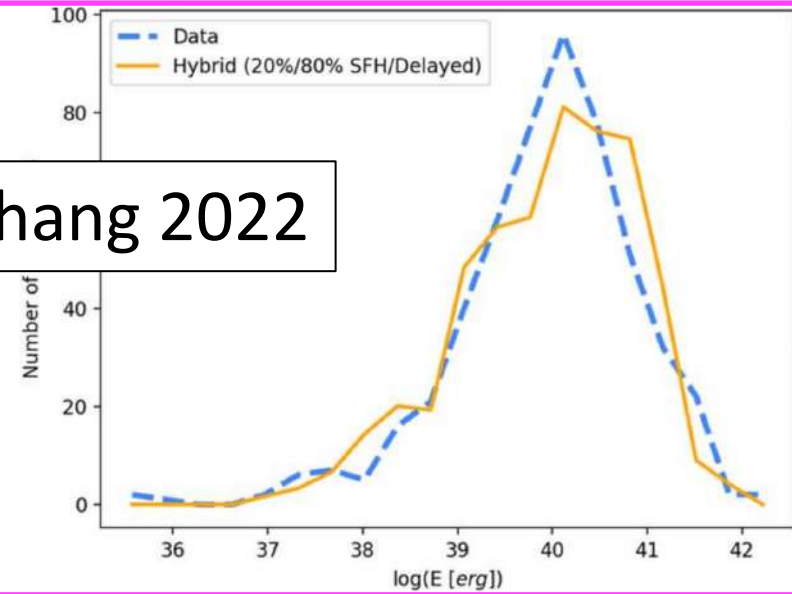
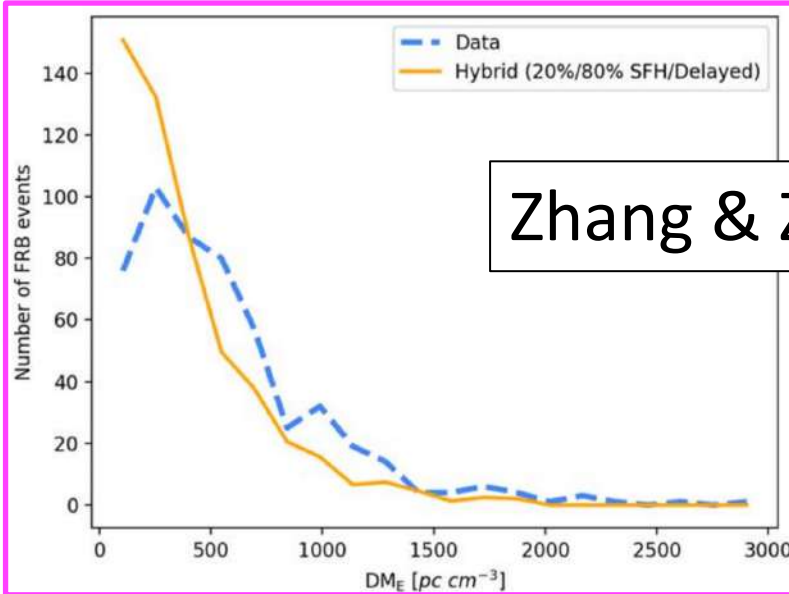


Stellar mass



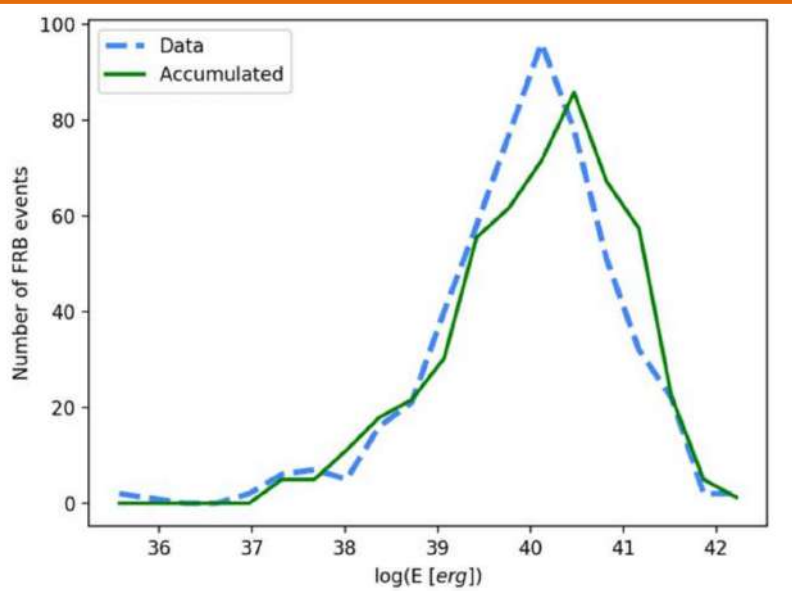
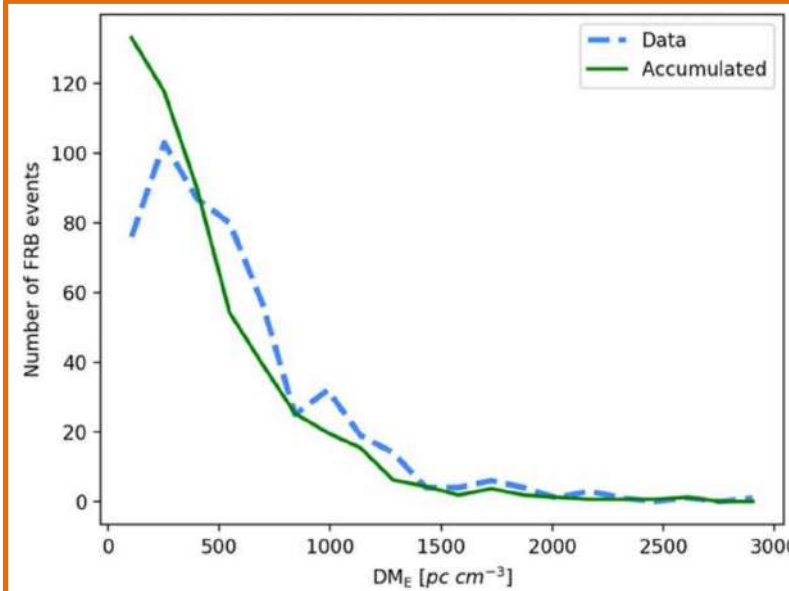
# Origin: including the old population better explains

Hybrid



Zhang & Zhang 2022

Stellar mass



# Conclusion

TH+2022a



x10 more samples than before

Non-repeater → Old objects

White dwarf



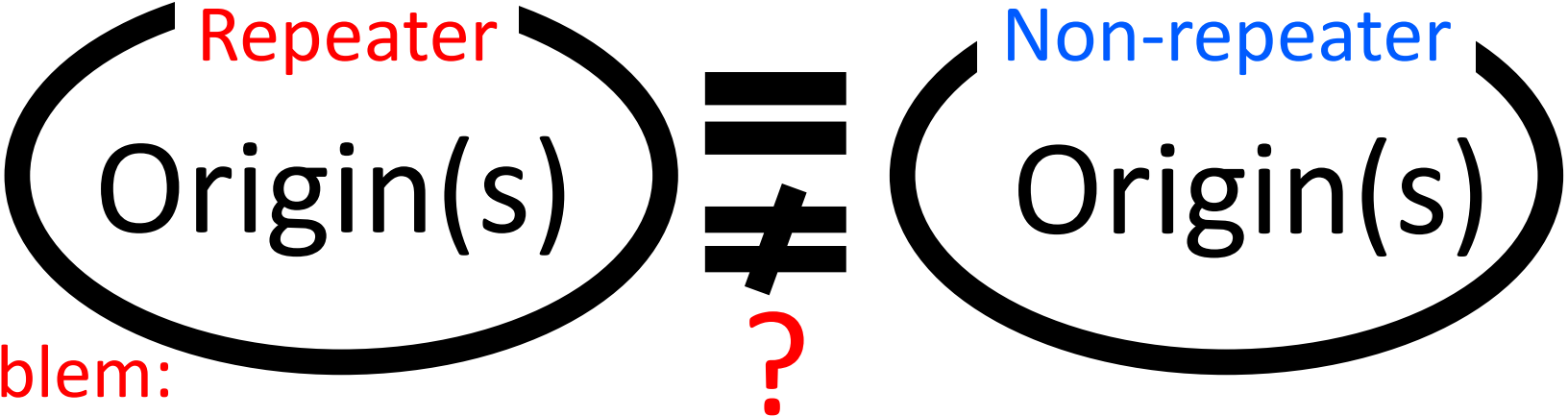
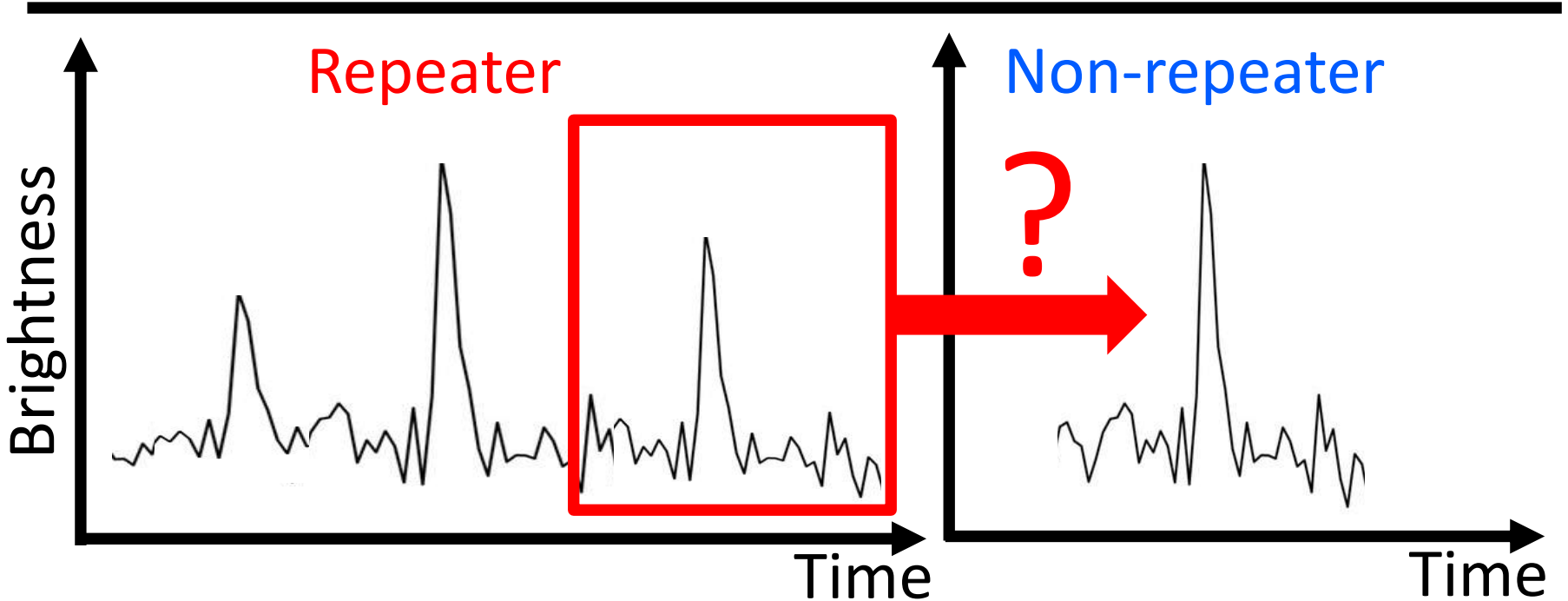
Old neutron star



Old stellar-mass black hole (BH)



# Origin: hard to classify Repeater/Non-repeater



Problem:  
need long monitoring observations with a high sensitivity

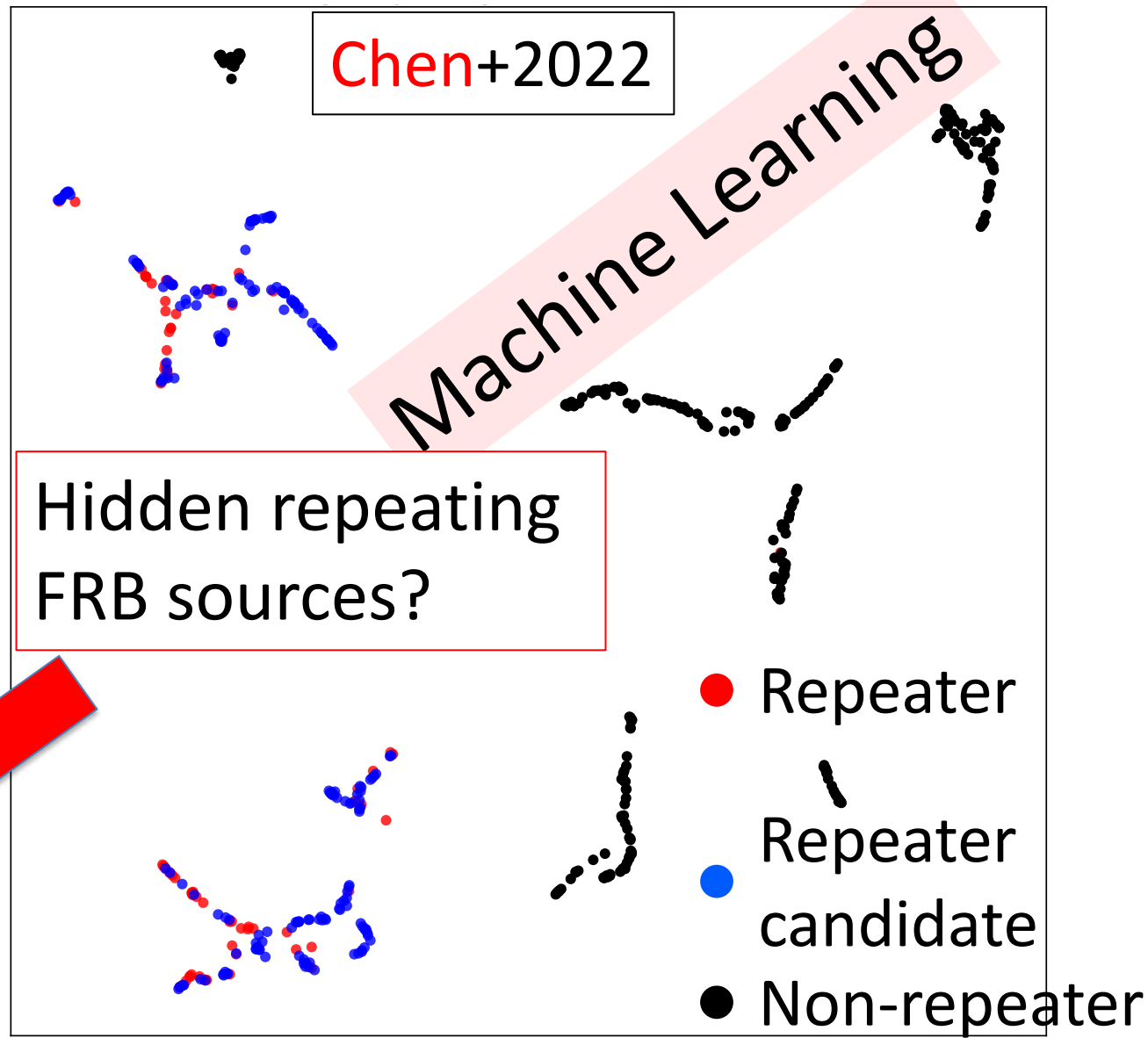


Oliver  
(Bo Han Chen)

testing w/ FAST



# ML may resolve it?





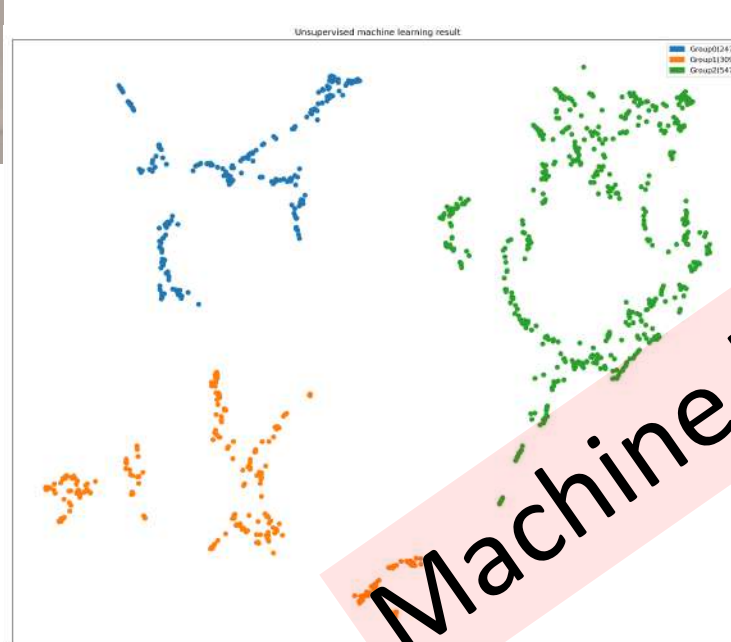
Oliver  
(Bo Han Chen)

# ML classification of repeating FRBs

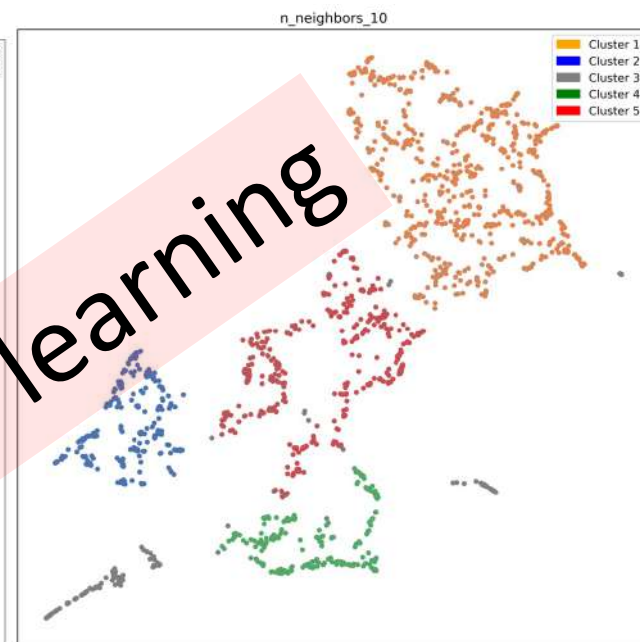


Chen+ in prep.

Jasper+ in prep.



FRB20201124 source



FRB121102 source



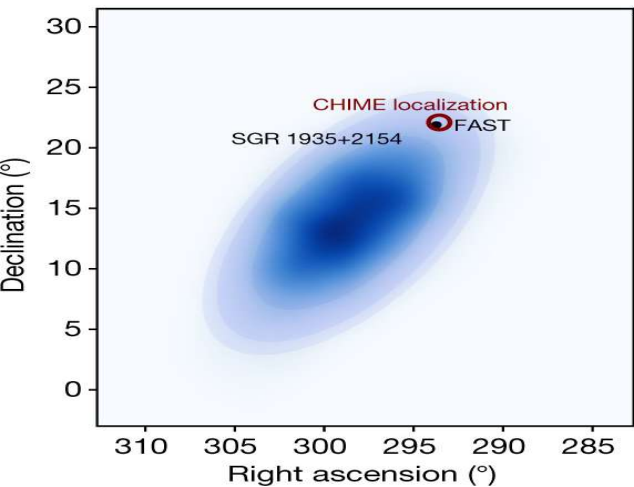
Jasper@RTU

Different physical mechanisms for each FRB source?



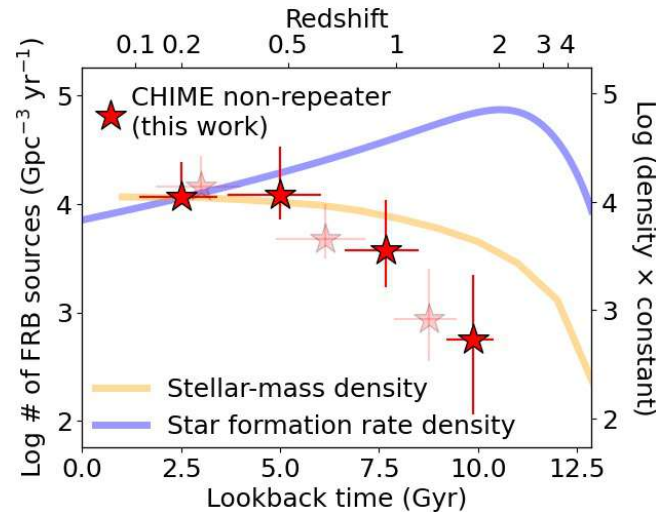
# Origin of FRBs: summary

e.g. Bochenek+2020



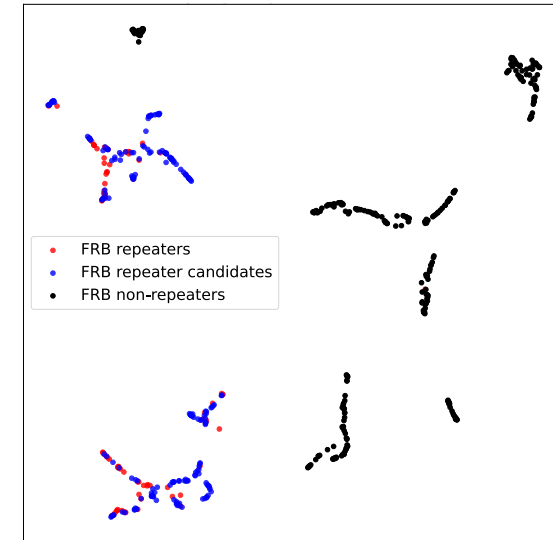
One confirmed case of a Galactic magnetar (repeater)

e.g. TH+2022



The number of non-repeaters  
↓  
old population

e.g. Chen+2022



ML approaches became feasible due to CHIME and FAST

# ***3. Applications of FRBs***

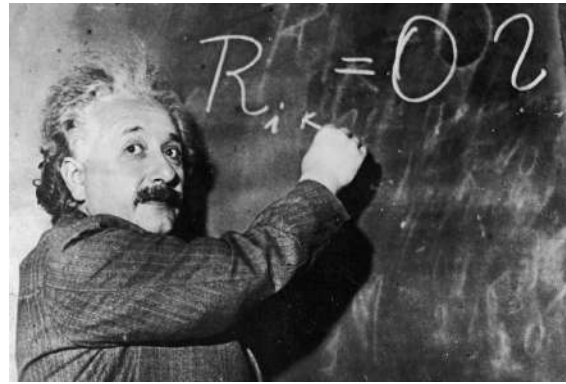
# **App:** key sciences to be addressed by FRBs

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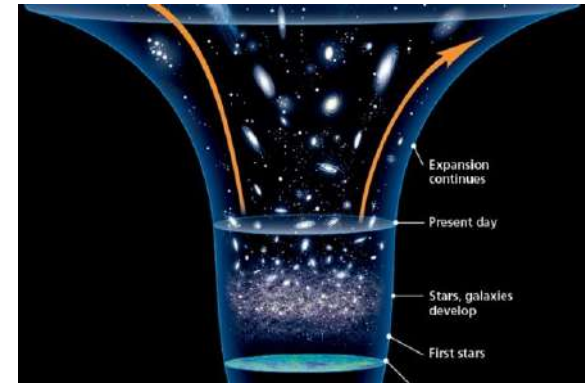
Missing baryon problem



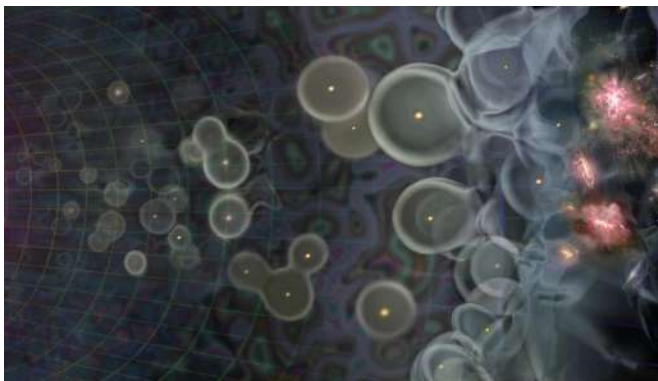
General relativity



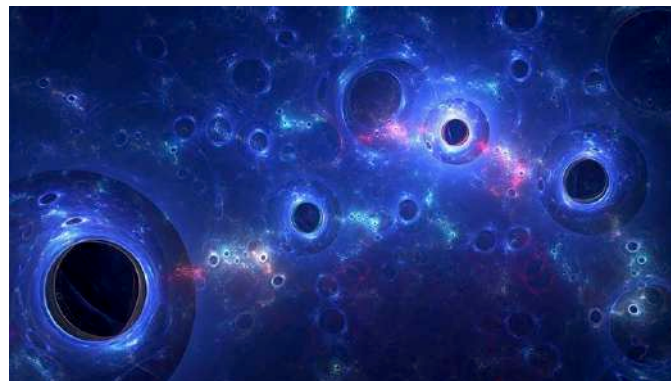
Dark energy



Cosmic reionization



Dark matter

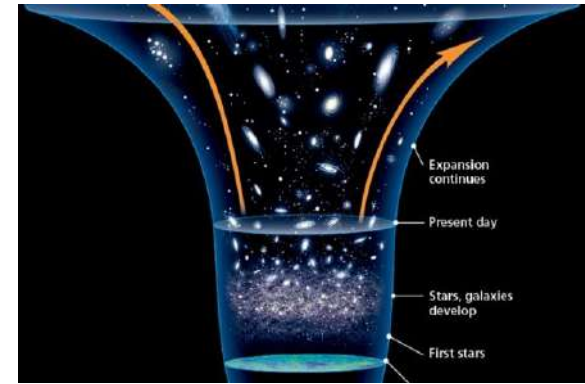
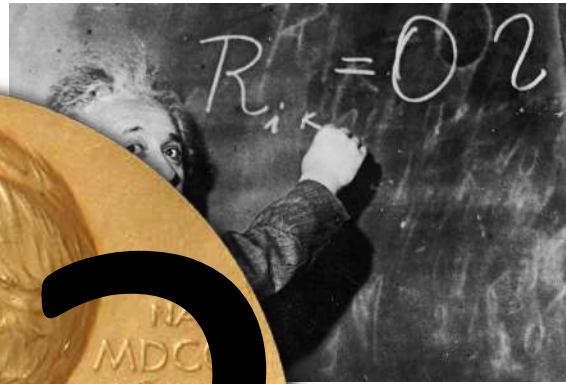


# **App:** key sciences to be addressed by FRBs

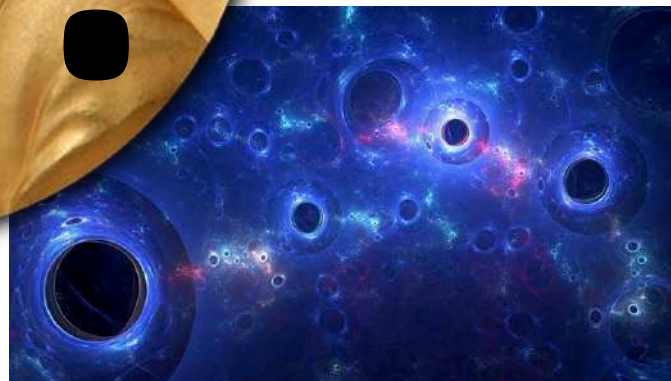
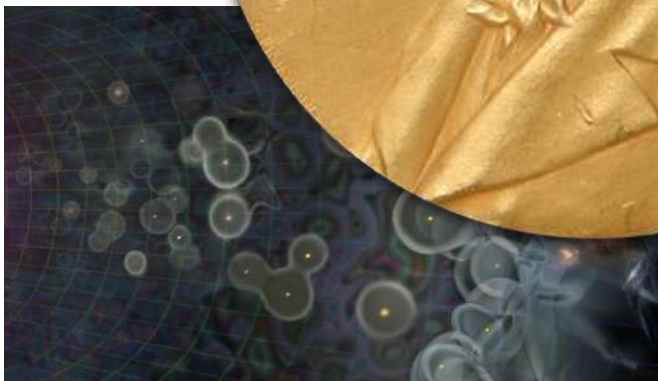
Missing baryon problem

General relativity

Dark energy



Cosmic reionization dark matter



# **App:** key sciences to be addressed by FRBs

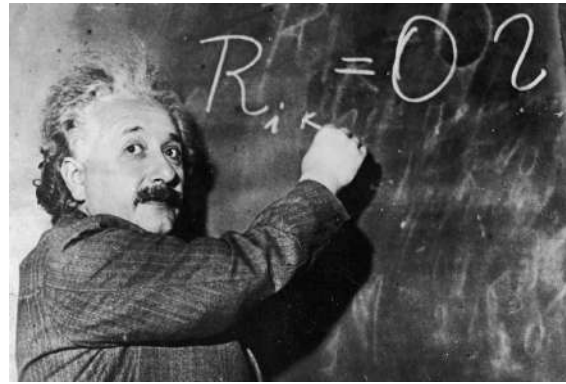
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Missing baryon problem



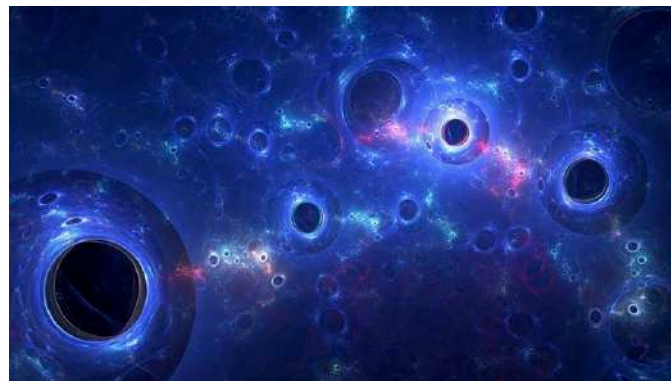
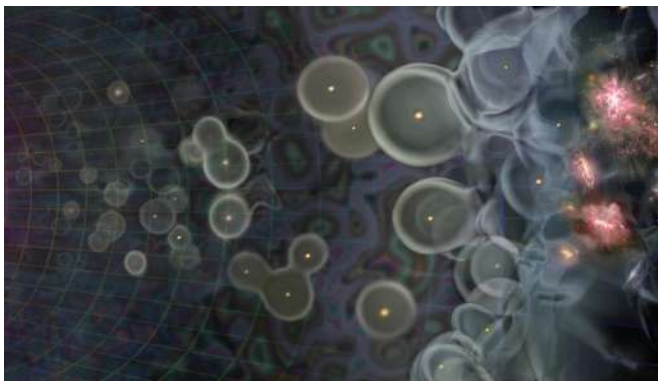
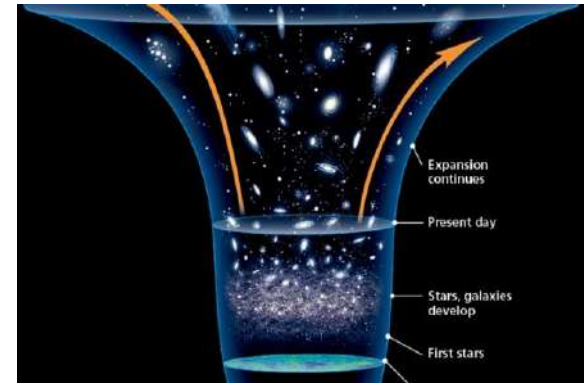
Cosmic reionization

General relativity



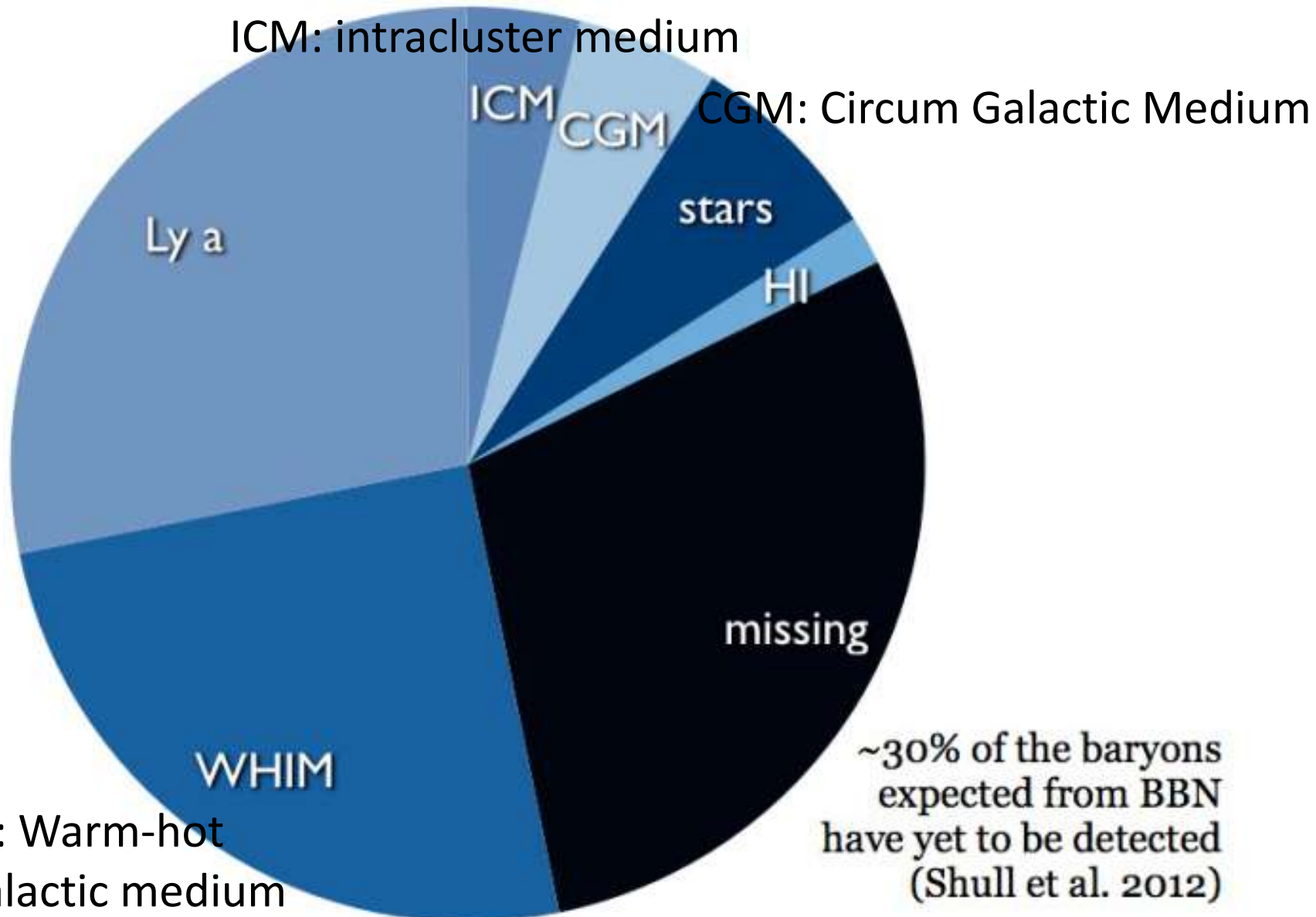
Dark matter

Dark energy



# *App:* the missing baryon problem

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# **Intro:** the speed of light changes in plasma

Speeds of radio emissions

high frequency: fast

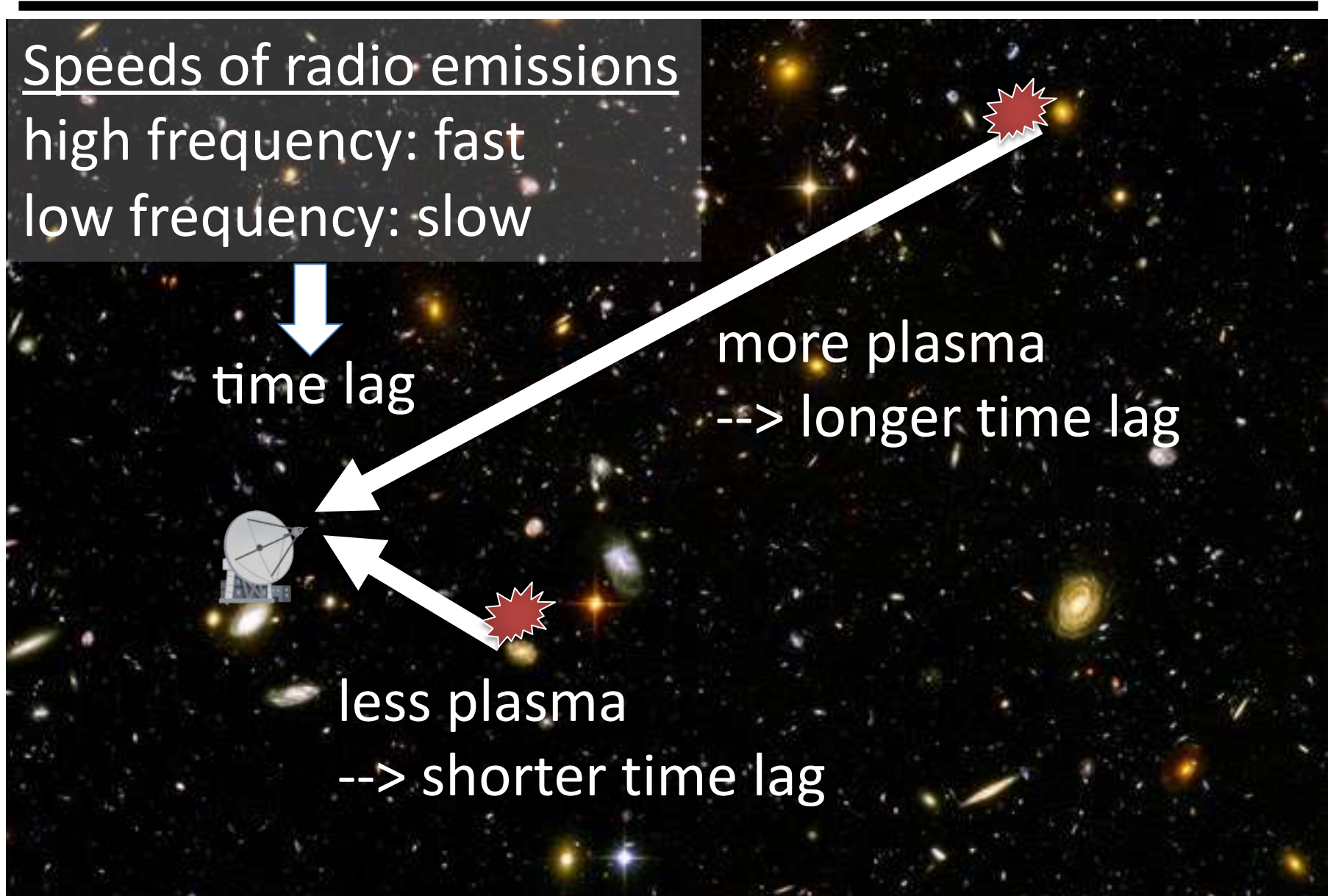
low frequency: slow

time lag

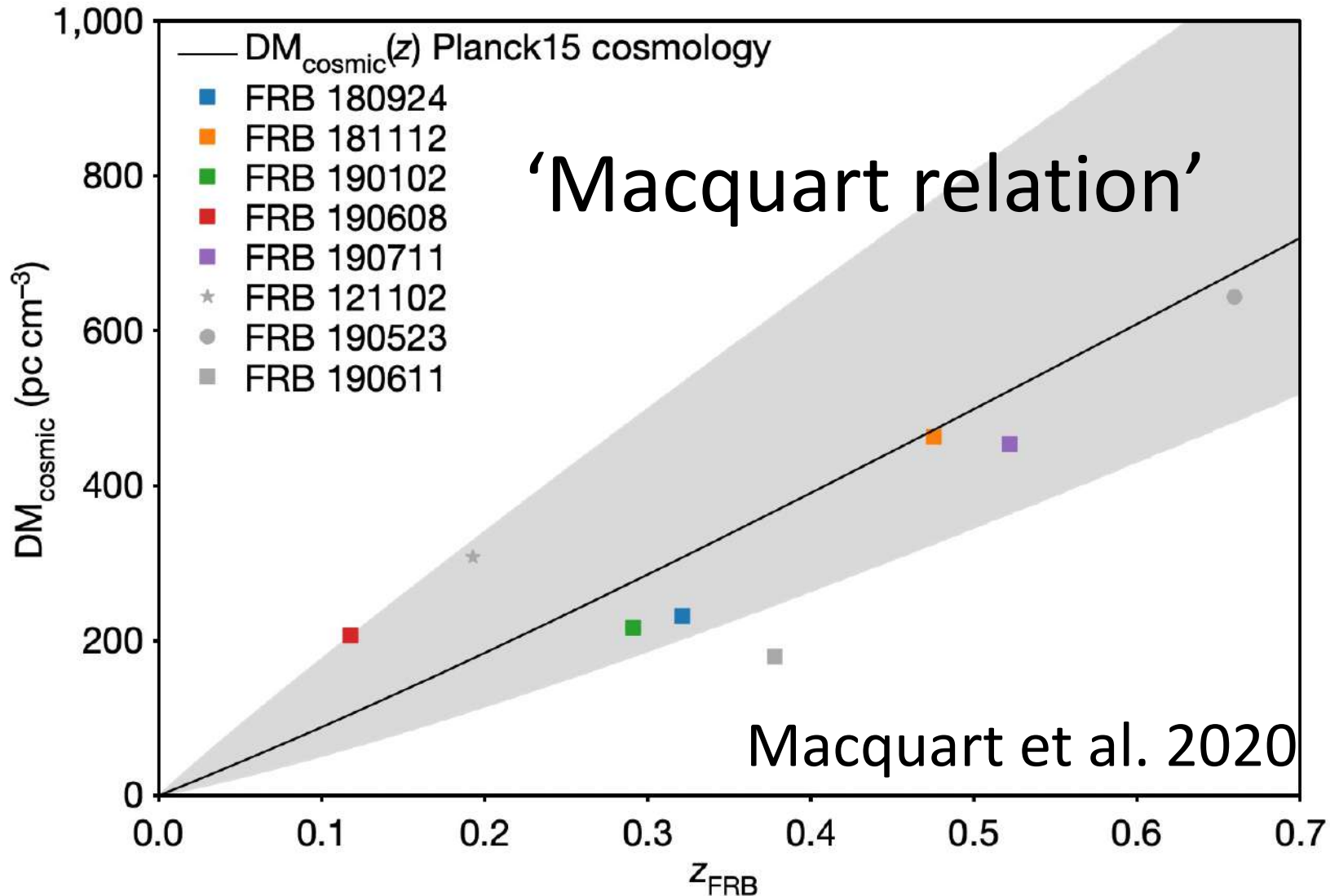


more plasma  
--> longer time lag

less plasma  
--> shorter time lag



# *App*: the missing baryon problem



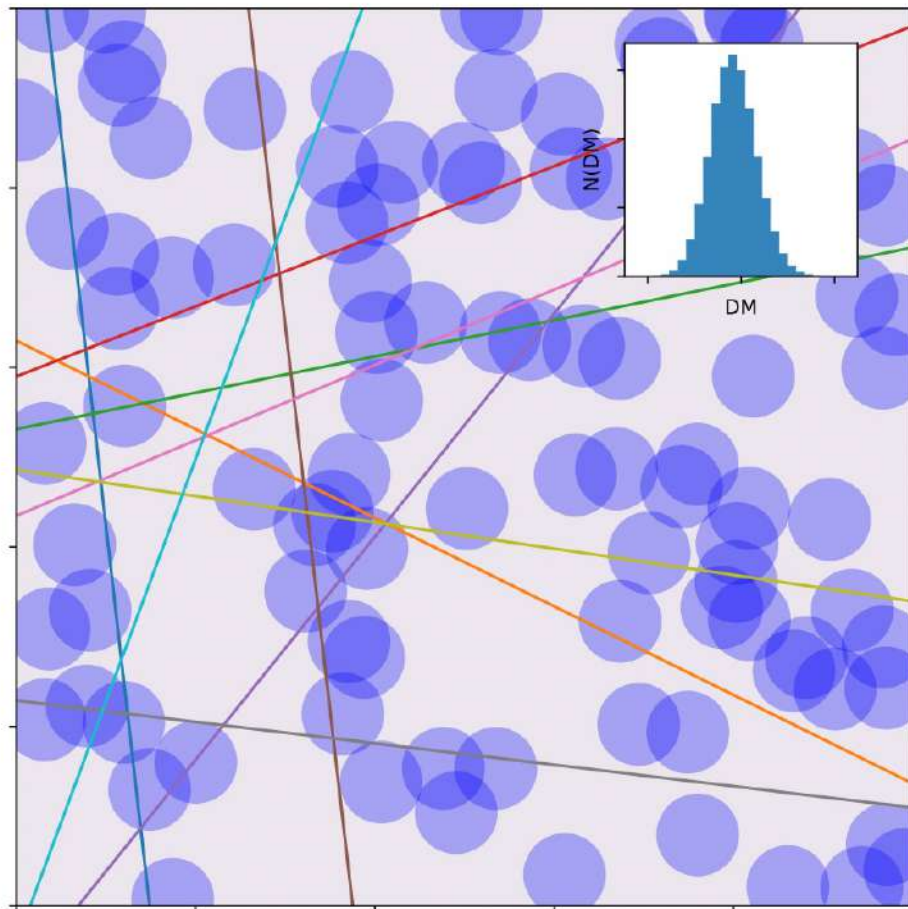
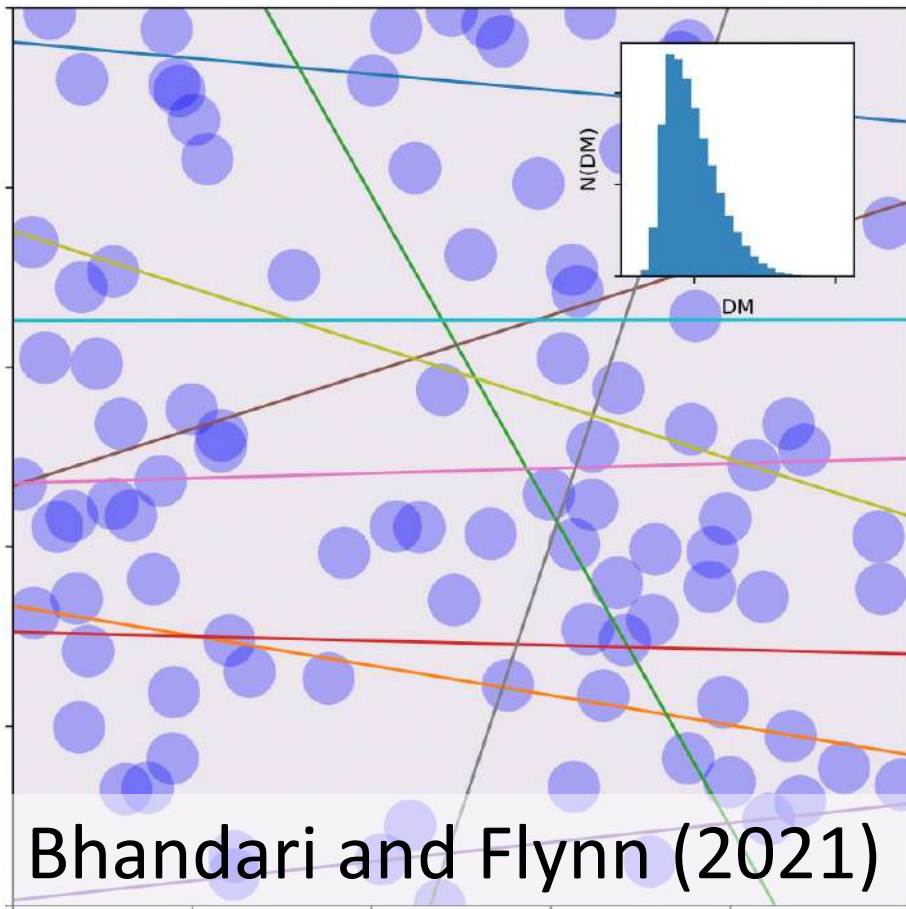
The next step would be the spatial distribution of baryon



# **App:** $\sim 100$ localized FRBs may reveal the baryon distribution

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## Predictions of dispersion measure distributions at a fixed redshift



***4. A new telescope plan  
in Taiwan: BURSTT***

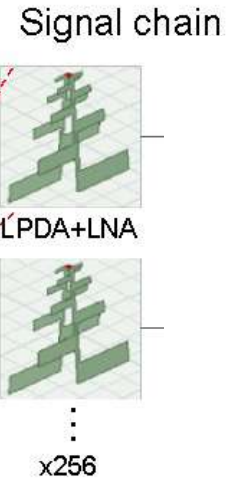
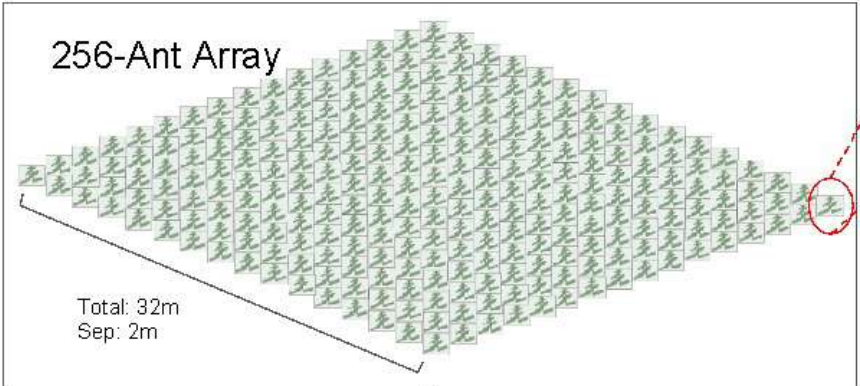
## ***BURSTT***: what are the bottlenecks of observations?

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- lack of localization capability
  - e.g., CHIME  $\sim$  arcmin
  - spec-z, host galaxy, progenitor etc.
- small FoV and low cadence
  - e.g., CHIME: 5-10 min per day (<1% of the day)
  - missing population of FRBs?
  - expensive for follow-up telescopes
- mismatched distance
  - GWs, neutrinos, high-time resolution follow-up
  - > nearby Universe

# Future FRB telescope in Taiwan

## Bustling Universe Radio Survey Telescope in Taiwan (BURSTT)

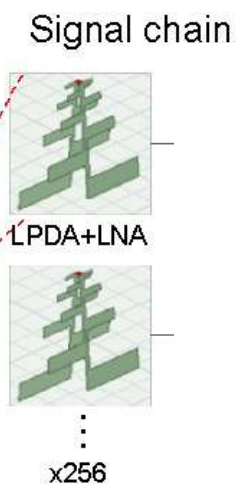
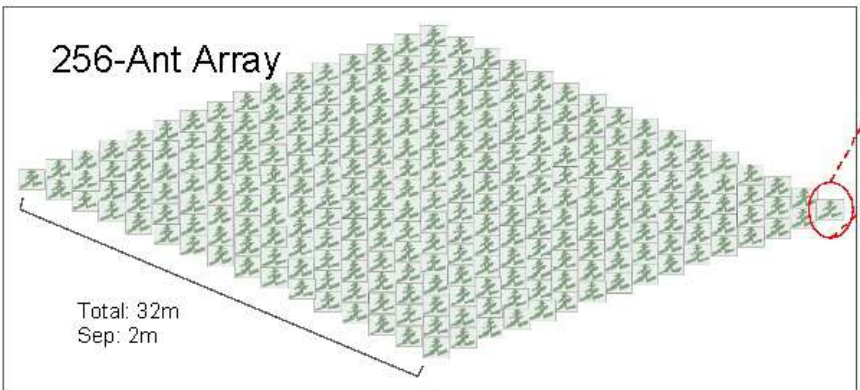


led by  
Prof. Pen, Ue-Li



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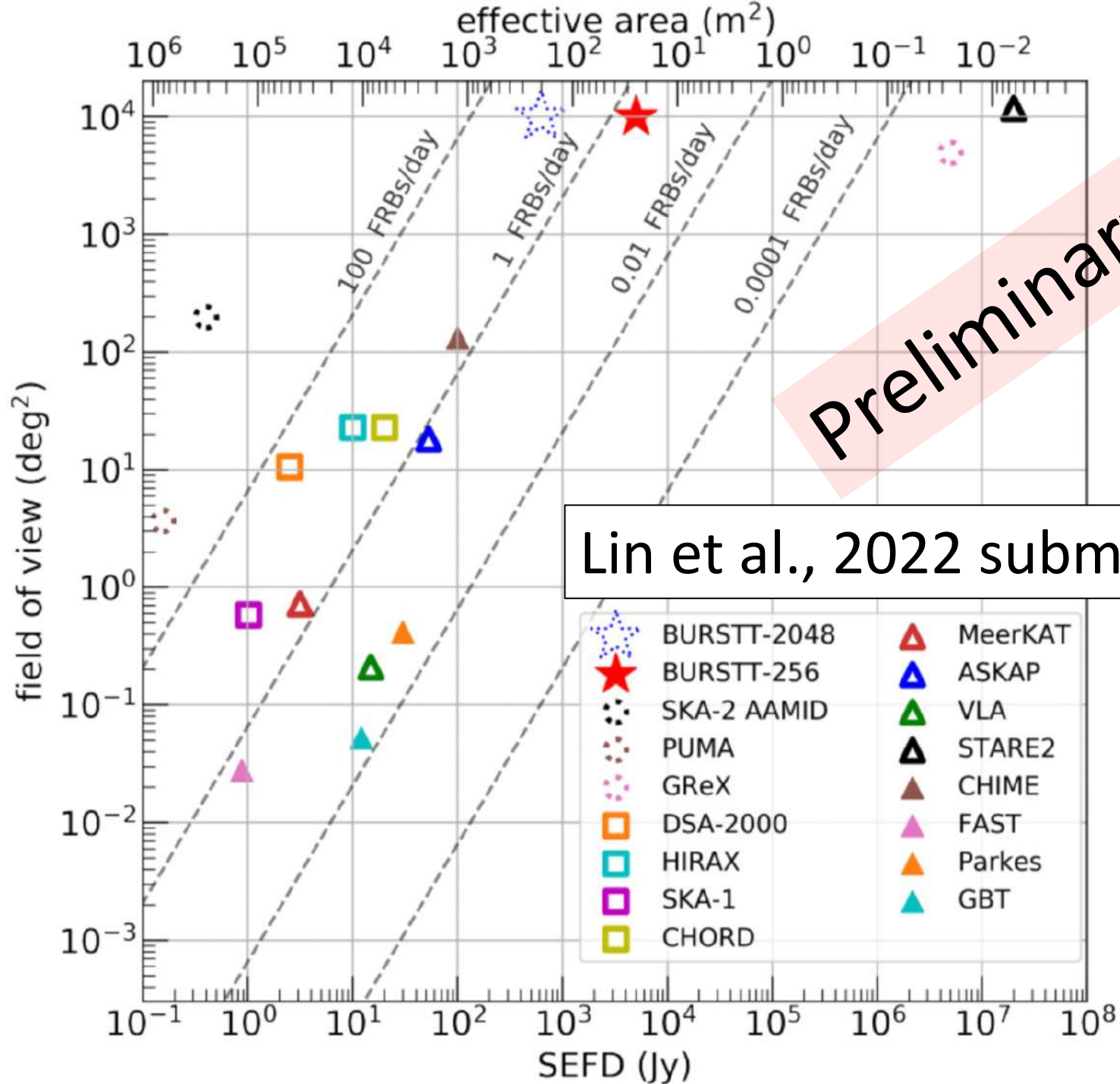


The world-best telescope to detect nearby FRBs

# ***BURSTT will resolve the bottlenecks***

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- lack of localization capability  
**BURSTT: sub-arcsecond**  
e.g., CHIME  $\sim$  arcmin
- small FoV and low cadence  
**BURSTT: 25 times better than that of CHIME**  
e.g., CHIME: 5-10 min per day (<1% of the day)
- mismatch in distance  
**BURSTT: dedicated to the nearby Universe**  
GWs, neutrinos, high-time resolution follow-up  
--> nearby Universe



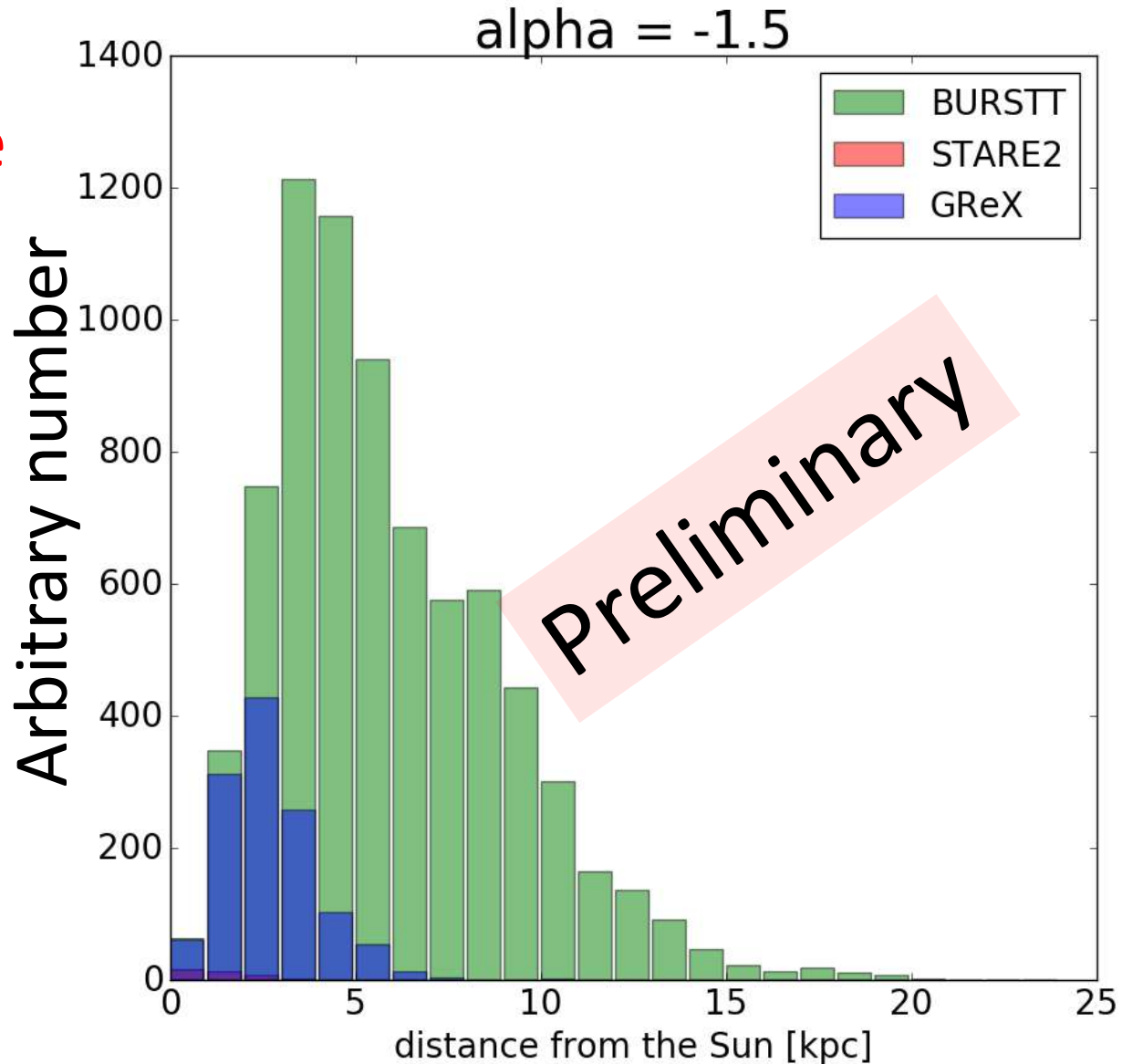
# Predictions of BURSTT FRBs

Number prediction  
of **Galactic FRB-like**  
**events** ( $E > 10^{32}$  erg)

cf. FRB from SGR  
1935+2154:  $10^{34-35}$   
erg



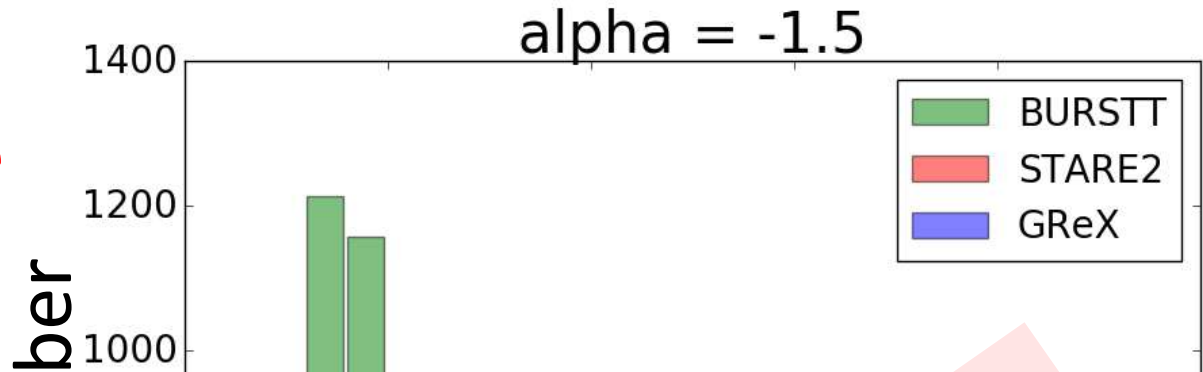
Decmend Lin





# Predictions of BURSTT FRBs

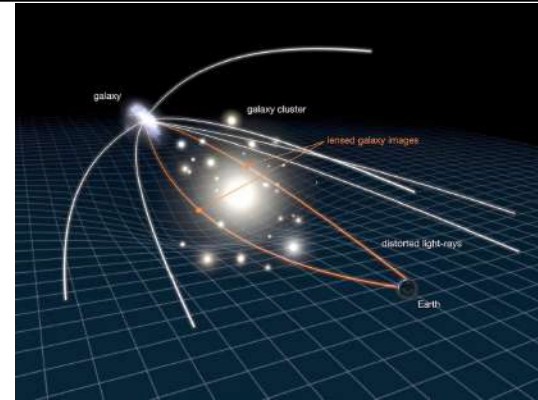
Number prediction of Galactic FRB-like events ( $E > 10^{32}$  erg)



System parameters	Ratio of value compared with STARE2	Ratio of detection rate
Area of antenna	× 56 larger	10
SEFD	× 20000 more sensitive	28
Central frequency	× 2 smaller	0.7
Bandwidth	× 2 larger	1.6
Field of view (FoV)	× 1.4 larger	5
Expected detection rate		156.8

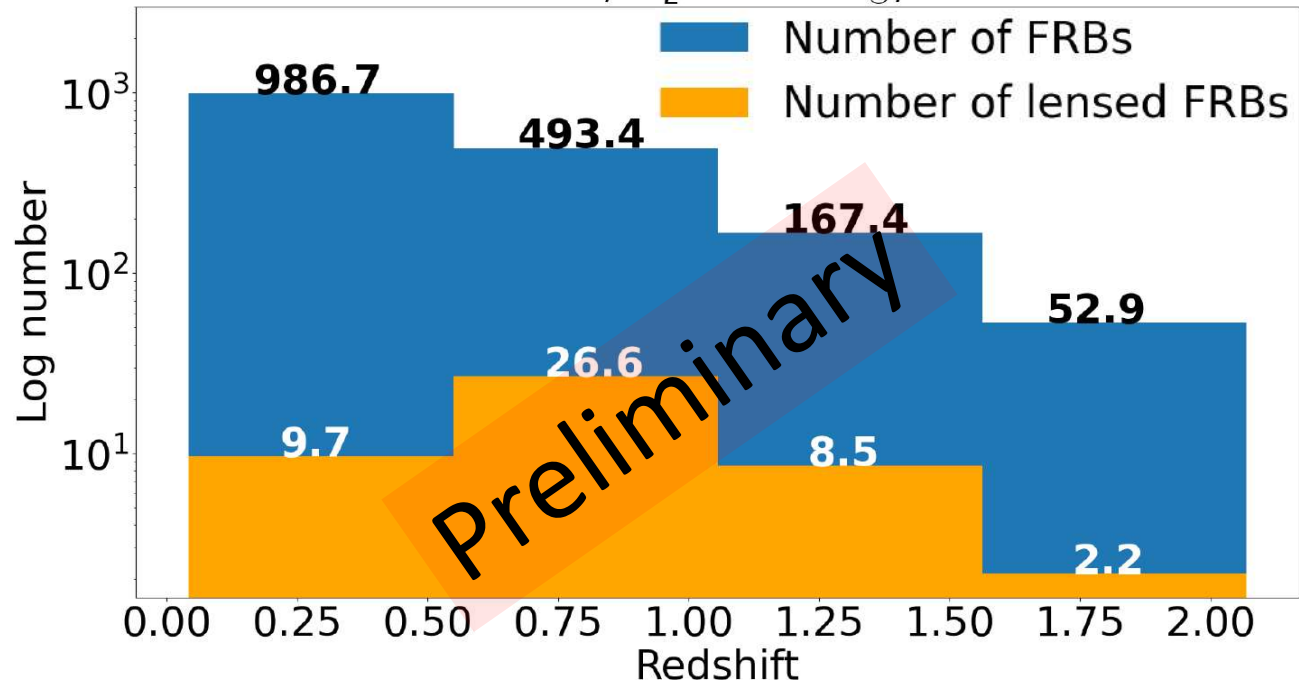
# Predictions of BURSTT FRBs

Number prediction of gravitationally lensed FRBs



Simon C.-C. Ho

BURSTT-2048,  $M_L = 0.001 M_\odot$ ,  $\Delta t = 10^{-9} \text{s}$



Preliminary

to constrain dark matter and cosmological parameters

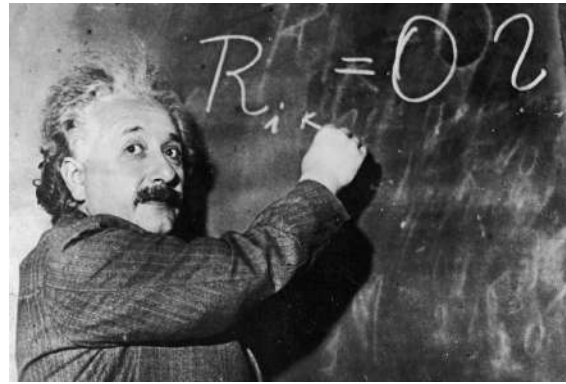
# ***Conclusion:*** FRB science is exciting

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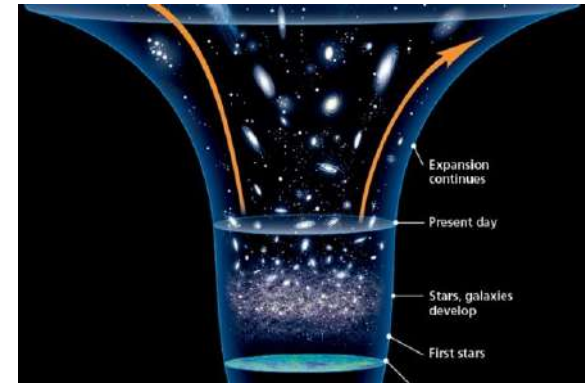
Missing baryon problem



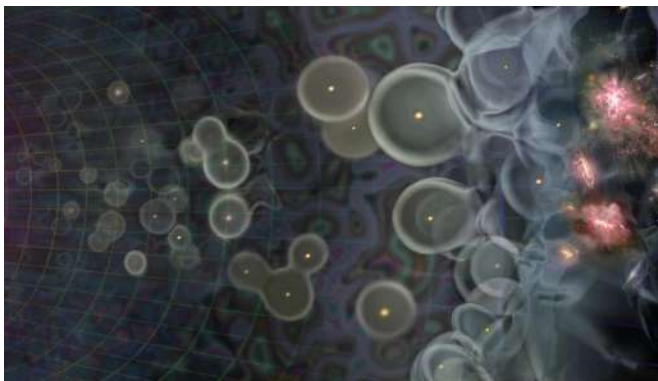
General relativity



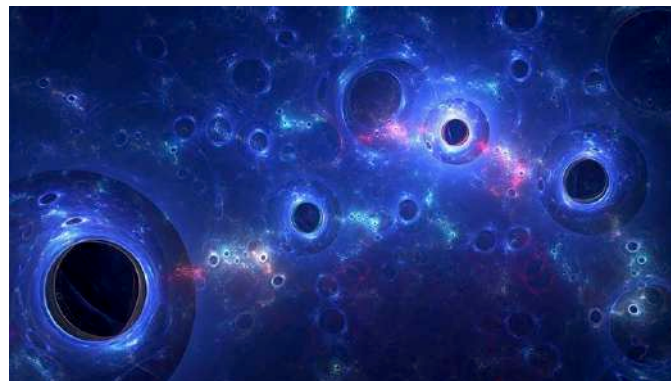
Dark energy



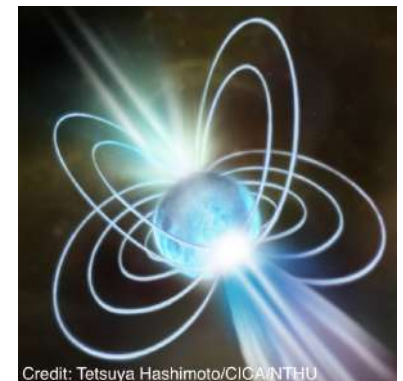
Cosmic reionization



Dark matter



The origin of FRBs



Credit: Tetsuya Hashimoto/CICANTHU

Backup slides



# Our result:

## do CHIME FRBs show the correlation?



Yes, at  $z < 0.3$

Kim et al. 2022

Seong Jin Kim

